WRAP UP OF TOPIC #5...

ELECTROMANGETIC RADAITAION & THE ELECTROMAGNETIC SPECTRUM

ATOMS vs MOLECULES

Quantum leap of electrons <u>WITHIN an ATOM</u> when photons are absorbed or emitted →



Quantum MOLECULAR MOTION behavior when photons are absorbed or emitted







Review p 26

Longwaves (LW)

The Electromagnetic Spectrum (another view)

> Shortwaves (SW)



Details on the spectrum: UV Vis IR

Visible Light range = 0.4 – 0.7 micrometers

Type of Electromagnetic Radiation	Range of Wavelengths (in units indicated)	Additional Information
Gamma rays	10⁻¹⁶ to 10⁻¹¹ in meters (m) using scientific notation	Involve high-energy processes <u>within a nucleus</u> caused by the strong force
UV Ultraviolet radiation UVC.2029 UVB.2932 UVA.3245	.0001 to 0.4 in micrometers (μm)	Involve electrons moving (quantum leaps)
VIS Visible light Solar	0.4 to 0.7 in micrometers (μm)	
IR Infrared radiation	0.7 to ~30 (up to 1000) in micrometers (μm)	Involve chaotic thermal <u>kinetic motion</u> of
IR Near Infrared radiation Longwave	0.7 - 1.0 in micrometers (μm)	IR photon
IR Far Infrared	1.0 - ~30 (up to 1000) in micrometers (μm)	Faster rotation rate
Microwaves	10 ⁻⁴ to 10 ⁻² in meters (m) using scientific notation	occur in nature & also electronically produced by a "magnetron" in a microwave oven
AM Radio waves	10 to 10 ² in meters (m) using scientific notation	occur in nature & also electronically produced in human-made electrical circuits

THE GREENHOUSE EFFECT

p

What is the relationship between . . . ENERGY E and FREQUENCY V OF PHOTONS?

KEY CONCEPT #1:

The Energy E of photons is directly proportional to their frequency v

 ∞ = "is proportional to"

 $\mathsf{E} \propto \mathsf{V}$

What is the relationship between . . . **ENERGY E** and **WAVELENGTH** λ OF PHOTONS?

KEY CONCEPT #2:

The Energy E of photons is inversely proportional to their wavelength λ E \propto C / λ f c = speed of light = a constant

SOLAR RADIATION: greatest intensity in SHORT wavelengths

(high energy & frequency)



EARTH **RADIATION:** entirely in LONG wavelengths (low energy & frequency)

The wavelength determines how the electromagnetic ENERGY (photon) will interact with MATTER !

NITROGEN GAS MOLECULE N₂





Rotation

Vibration



Slow rotation rate

H H Faster rotation rate

CARBON DIOXIDE GAS MOLECULE CO₂





So what is a Greenhouse Gas?

abbreviation we'll use = GHG

GHG = a gas than can absorb and emit (re-radiate) <u>INFRARED</u> wavelengths of Electromagnetic Radiation



> 0.7 - 1000 micrometers





The QUANTUM BEHAVIOR of **CERTAIN MOLECULES** with respect to **INFRARED RADIATION** is the **REASON THAT GREENHOUSE GASES ARE GREENHOUSE GASES!!**

TOPIC # 6 The RADIATION LAWS The KEY to unlocking the topics of: The GREENHOUSE EFFECT, GLOBAL WARMING & OZONE DEPLETION!



TODAY: Topic #6 – Part I pp 29-31

$E = \sigma T^{A}$ "The equations we seek are the poetry of nature Why is nature that way?



Why is it possible for these powerful manifestations of forces to be trapped in a very simple, beautiful formula?"

"This has been a question which many people have discussed, but there is no answer."

E=hc/2

~ Chen Ning Yang (b. 1922) US physicist





FIRE UP YOUR CLICKERS FOR RESPONSE TIME!

Clicker



ResponseWare Device

Open your "APP" or login to: <u>rwpoll.com</u>

ENTER CHANNEL = 28

ENTER SESSION ID (see above)

A "cartoon" view of Solar vs Terrestrial radiation:

Both Sun & Earth are radiating energy . . .



(Sun & Earth are NOT shown in proper scale!)



Warm up question: Which one do you think is the <u>most</u> <u>accurate</u> depiction of the Greenhouse Effect??



= gases in the atmosphere

= solar (shortwave) radiation (High Energy) = terrestrial (longwave) radiation (Lower Energy)

Warm up question: Which one do you think is the <u>most</u> <u>accurate</u> depiction of the Greenhouse Effect??



= gases in the atmosphere

= solar (shortwave) radiation (High Energy) = terrestrial (longwave) radiation (Lower Energy)



Actually, NONE of these is <u>exactly</u> correct, and we will learn why in a future lecture.... but for now, **B is the preferred answer** see the image on bottom of p 27 in Class Notes.

REVIEW



KEY BANDS IN THE SPECTRUM FOR GLOBAL CHANGE: UV, Visible, IR, NIR,

Type of Electromagnetic Radiation	Range of Wavelengths (in units indicated)	Typical Source
Gamma rays	10^{-16} to 10^{-11}	high-energy processes within nucleus caused
Ultraviolet radiation Sol Visible light SV	ar .0001 to 0.4 in micrometers (µm) 0.4 to 0.7	electrons moving (quantum leaps) within individual atoms
Infrared radiation Near Infrared radiation Far Infrared	Terrestrial LW (IR)	chaotic thermal kinetic motion of molecules due to their thermal energy IB photon Faster rotation rate
Microwaves	10 ⁻⁴ to 10 ⁻² in meters (m) using scientific notation	electronically produced by microwave oven
AM Radio waves	10 to 10 2 in meters (m) using scientific notation	electronically produced waves vibrate in human-made electrical circuits

Review p 28



THE RADIATION LAWS !!!

Keys to Understanding the Greenhouse Effect





Emission of radiation

<u>All</u> substances emit radiation as long as their temperature is above absolute zero

(-273.15°C or 0 Kelvin).



TYING THE RADIATION "LAWS" to GLOBAL CHANGE

The Sun's energy (solar) is emitted in the form of electromagnetic radiation.

mostly Shortwave (SW) (but also some LW) The Earth's energy (terrestrial) is also emitted in the form of electromagnetic wavelengths...

but in LW INFRARED (IR) wavelengths



LAW #2 BLACKBODY & PLANCK FUNCTION CONCEPT

(NOTE: the Earth is <u>NOT</u> a VERY good "black body" emitter)

Black body (def): a body that emits equally well at all wavelengths (i.e. radiates with 100% efficiency)

It also absorbs equally well at all wavelengths and is a "perfect absorber" (hence described as "black") Blackbodies ("ideal emitters") exhibit a *defined relationship* between:

the <u>intensity</u> of radiation <u>energy</u> (E) (i.e. amount of radiation flux) they give off & the <u>wavelength</u> of that radiation.

This relationship is called the **Planck function**:

E = h * speed of light / wavelength

Where (h) is Planck's constant.

<mark>E = h c / λ</mark>

Or

Planck Function:

The Sun emits energy at ALL wavelengths...

but the amount of Energy emitted is inversely related to the wavelength of emission "I radiate at the speed of light like a blackbody; but my energy flux is GREATEST at SHORTER wavelengths"



This can be depicted in a graph:



An emitting blackbody's <u>SHORTER</u> wavelengths have HIGHER intensity radiation (and greater energy flux) than the LONGER wavelengths Easy way to remember the PLANCK FUNCTION / BLACKBODY concept:

"The shorter the <u>wavelength</u>, the GREATER the intensity of the <u>energy flux</u>"

Q1 - Gamma radiation involves a greater energy flux than microwave radiation.

- 1. True
- 2. False



- 3. Both wavelength bands have the same energy flux
- 4. We haven't learned enough yet to answer this!

Q1 - Gamma radiation involves a greater energy flux than microwave radiation.





- 3. Both wavelength bands have the same energy flux
- 4. We haven't learned enough yet to answer this!

LAW #3: THE STEFAN-BOLTZMANN LAW:

If the substance is an ideal emitter (black body),

The total AMOUNT of radiation given off is proportional to the fourth power of its absolute TEMPERATURE.



where σ is a constant (the Stefan-Boltzmann constant) which has a value of 5.67 x 10 ⁻⁸ W/m⁻² (or 5.67 x 10 ⁻⁸ J/m²) and *T* is the absolute temperature (in Kelvin)

Energy =
$$\sigma T^4$$



Stefan-Boltzmann Law (easy way) This law links:

(E) the <u>total</u> amount of <u>energy flux</u> that is emitted by a blackbody & (T) the body's <u>temperature</u>

(specifically, the <u>4th power</u> of the body's absolute temperature)

"the hotter the body, the (much) greater the amount of energy flux or radiation"



The total amount of energy flux described by the Stefan-Boltzmann Law is proportional to the <u>area under</u> the Planck function curve





Figure on p 42 in SGC E-text

Stefan-Boltzmann Law:

"I'm HOT, so I emit LARGE amounts of high intensity energy"



"I'm COOL, so I emit much LESS energy . . . AND <u>MY</u> energy is at a lower intensity than Mr. Sun over there!"



Why is this concept important? Because it means that:

> the <u>amount</u> of radiation given off by a body is a <u>very *sensitive*</u> function of its <u>temperature</u>

Therefore . . . small changes in temperature can lead to BIG changes in the amount of radiation given off.





Q2 – Which would you use: the <u>Planck Function</u> or the <u>Stefan-Boltzmann Law</u> to accurately compute <u>the total amount of ENERGY</u> emitted to space by planet Earth?

- 1. The Planck Function
- 2. The Stefan Boltzmann Law
- 3. Both of them together
- 4. Neither one is appropriate because the Earth is NOT a blackbody



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- **1. The Planck Function**
- 2. The Stefan Boltzmann Law
- 3. Both of them together

4. Neither one is appropriate because the Earth is NOT a blackbody

Q3 – Which would you use: the Planck Function or the Stefan-Boltzmann Law to compute the total amount of energy emitted to space by planet Earth, IF you assume the Earth emits like a blackbody & you know the Earth's temperature?

- **1. The Planck Function**
- 2. The Stefan Boltzmann Law
- 3. Neither one is appropriate because you would need to know the wavelengths of radiation the Earth emits
- 4. Don't know



Q3 – Which would you use: the Planck Function or the Stefan-Boltzmann Law to compute the total amount of energy emitted to space by planet Earth, IF you assume the Earth emits like a blackbody & you know the Earth's temperature?

1. The Planck Function

2. The Stefan Boltzmann Law

3. Neither one is appropriate because you would need to know the wavelengths of radiation the Earth emits

4. Don't know

How to do it:

 $\mathbf{E} = \boldsymbol{\sigma} T^4$

E = Energy per unit area, so all we need to know is the AREA of the emitting Earth's surface + what T is.

From geometry: Do you remember the formula for computing the area of a sphere?

The area of a sphere of radius R is





 $\mathbf{E} = 4 \prod \mathbf{R}^2 \times \mathbf{\sigma} \mathbf{7}^4$

See box on p 44 in SGC E-text for more details

10 minute SUSTAINABILITY SEGMENT more of:



Then ... MORE CLICKER PARTICIPATION POINTS Breaking news on gas and coalfired plant regulations!

EPA to revise climate rule for new power plants; will still require carbon capture

By Juliet Eilperin, Published: September 11 E-mail the writer 🦘

This month, the <u>Environmental Protection Agency</u> will propose standards that will establish stricter pollution limits for gas-fired power plants than coal-fired power plants, according to individuals who were briefed on the matter but asked not to be identified because the rule was not public yet.

The revised rule, which would impose <u>greenhouse gas limits on power plants</u> for the first time, will still require utilities to install costly carbon controls on coal plants.

More health and science news New EPA climate proposal still tough on coal



Juliet Eilperin SEP 11 EPA will unveil a proposal this month to limit carbon emissions from new coal and gas plants, sources say.

Today's news in The Washington Post, The Wall Street Journal, and others . . . The agency initially proposed that any new power plant emit no more than 1,000 pounds of carbon dioxide per megawatt hour of electricity produced but <u>decided to</u> <u>overhaul that rule</u> this spring out of concern that it would face a stiff legal challenge.

The average U.S. natural gas plant emits 800 to 850 pounds of carbon dioxide per megawatt, and coal plants emit an average of 1,768 pounds. According to those familiar with the new EPA proposal, the agency will keep the carbon limit for large natural gas plants at 1,000 pounds but relax it slightly for smaller gas plants. The standard for coal plants will be as high as 1,300 or 1,400 pounds per megawatt hour, the individuals

LAW # 4: Temperature and wavelength

As substances get HOTTER, the wavelength at which radiation is emitted will become SHORTER.

This is called Wien's law.

Wien's Law can be represented as: $\lambda_{m} = a/7$

where \mathcal{M}_{m} is the WAVELENGTH in the spectrum at which the energy peak occurs,

(m indicates "max")

7 is the absolute **TEMPERATURE** of the body, and

a is a constant (with a value of 2898)

(if λ_{m} is expressed in micrometers.)



Note the INVERSE relationship between wavelength and temperature

See figure on p 42 in SGC

Wien's Law (easy way)

λ max = constant / T

(Inverse relationship between wavelength and temperature)

"The <u>hotter</u> the body, the <u>shorter</u> the wavelength"

"The <u>cooler</u> the body, the <u>longer</u> the wavelength"

Wien's Law -- Why is this concept important?

Because it means that very HOT objects (like the sun) that radiate like blackbodies will radiate the maximum amount of energy at SHORT wavelengths,

while COOLER bodies will radiate most of their energy at LONGER wavelengths.



Wein's Law:

"I'm HOT, so I emit my maximum amount of radiation at SHORTER wavelengths"



"I'm COOL, so I emit my maximum amount of radiation at LONGER wavelengths"



LW = infrared (IR)

SW = visible (VIS) & ultraviolet (UV)



NOTE: this is a logarithmic scale -- values increase exponentially to the right

Another view of the same concept:



Know & understand what this figure is illustrating!

p 30

THE RADIATION LAWS Re-cap of Laws # 2 - 4

Planck Function:

 $E = h c / \lambda$

The Sun can emit energy at ALL wavelengths, but the amount of energy emitted is inversely related to its wavelength.

"I radiate at the speed of light like a blackbody; most of my energy is emitted at shorter wavelengths "

Planck function (blackbody radiation curve) Wavelength

Stefan-Boltzmann Law:

$E = \sigma T^4$

'I'm HOT, so I emit LARGE amounts of high intensity energy"



"I'm COOL, so I emit LESSER amounts of energy; *plus* my ENERGY is at a lower intensity than Mr. Hotshot over there!"



Wein's Law: 🤈

"I'm HOT, so I emit my maximum amount of radiation at SHORTER wavelengths"

= a / T

"I'm COOL, so I emit my maximum amount of radiation at LONGER wavelengths"



SW = visible & ultraviolet (UV)

LW = infrared (IR)



Q4 – Which choice correctly matches the Stefan-Boltzmann LAW with its "mantra" (A, B, C):

A "The <u>hotter</u> the body, the <u>shorter</u> the wavelength" The <u>cooler</u> the body, the <u>longer</u> the wavelength"

B "SHORTER wavelengths have HIGHER intensity radiation than LONGER wavelengths "

Top of p 31

C "The hotter the body, the (much) greater the amount of energy flux or radiation"







"The <u>hotter</u> the body, the <u>shorter</u> the wavelength" The <u>cooler</u> the body, the <u>longer</u> the wavelength"

(B) Planck Function:

"SHORTER wavelengths have HIGHER intensity radiation than LONGER wavelengths"

(C) Stefan-Boltzmann Law:

$$E = \sigma T^4$$

"The hotter the body, the (much) greater the amount of energy flux or radiation"

RADIATION LAWS # 5 and #6

to be continued

Postscript: Inside a Solar Cell!

From:

Starring:





http://www.pbs.org/wgbh/nova/tech/saved-by-the-sun.html



Is it time to take SOIAL ENELGY seriously?

le Sill



Phosphorus (P) "doped" Si layer



Extra (P) electrons move down to (B) layer, negatively charging it

Silicon (Si)

Boron (B) "doped" Si layer



Read this explanation at:

http://www.pbs.org/wgbh/nova/tech/how-solar-cell-works.html/



INSIDE A SOLAR CELL



5-277

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GO CATS!