

**Topic # 6 ATMOSPHERIC
STRUCTURE
&
CHEMICAL COMPOSITION
Part II**

Plus wrap up of
some other topics

Individual Test #1 Class Statistics

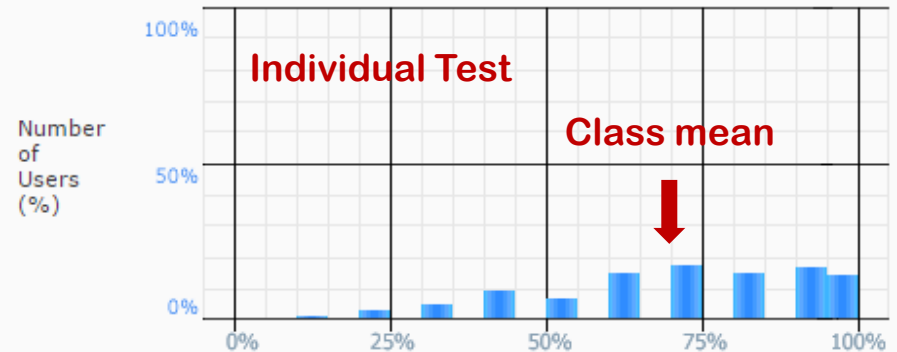
Number of submitted grades: 110 / 110



THE POWER OF COLLABORATIVE LEARNING!

REMEMBER:
Your group is there to **HELP YOU LEARN** by working out problems together that & helping each other understand concepts

Grade Distribution

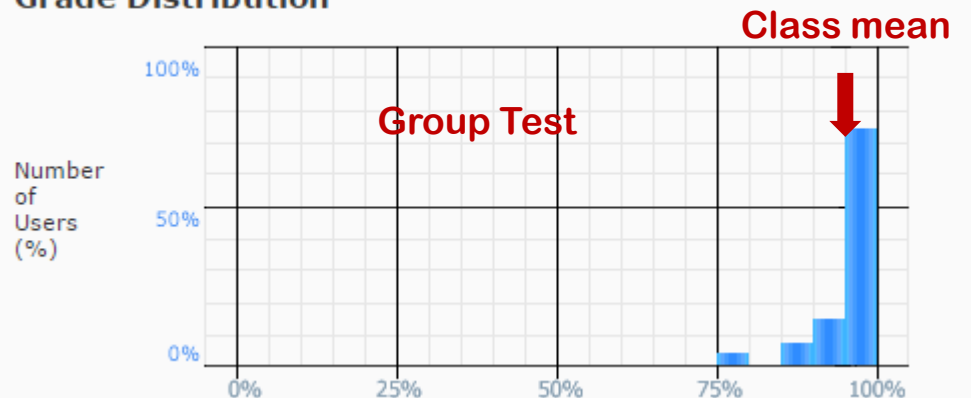


Group Test #1 Class Statistics

Number of submitted grades: 110 / 110



Grade Distribution



ABOUT THE LEARNING PHILOSOPHY UNDERLYING THIS COURSE

FAQ #2: <http://www.ltrr.arizona.edu/kkh/natsgc/faq.htm#2>.

This course is designed to help you learn the course material in ways that are different from traditional lecture courses. . . .

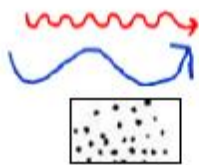
Class time is valuable: Class should be active w/ hands-on activities so that students can be engaged in their own learning

- Students learn best when they come to class prepared (by reading, Self tests &RQs) to which new material is added
- Students learn best when they have a common "starting place" for deepening their understanding (why Dr H reviews concepts)
- A group of people working together cooperatively can solve complex problems more easily and at a higher level than an individual working alone
- To achieve success in group activities, each individual has to be perceived as someone with something to offer, and each individual is responsible for preparing and contributing his or her own individual knowledge.

GROUP WORK SELF TEST

SOLAR vs TERRESTRIAL RADIATION CLASS CONCEPTS SELF TEST

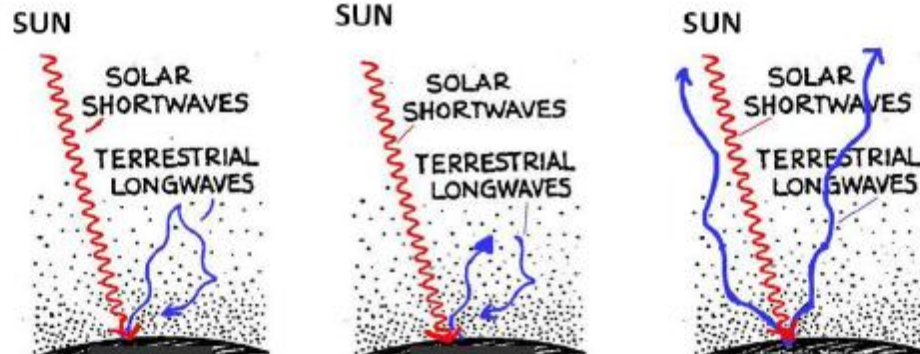
KEY:



= represents Solar shortwave (SW) radiation

= represents Terrestrial longwave (LW) (infrared IR radiation)

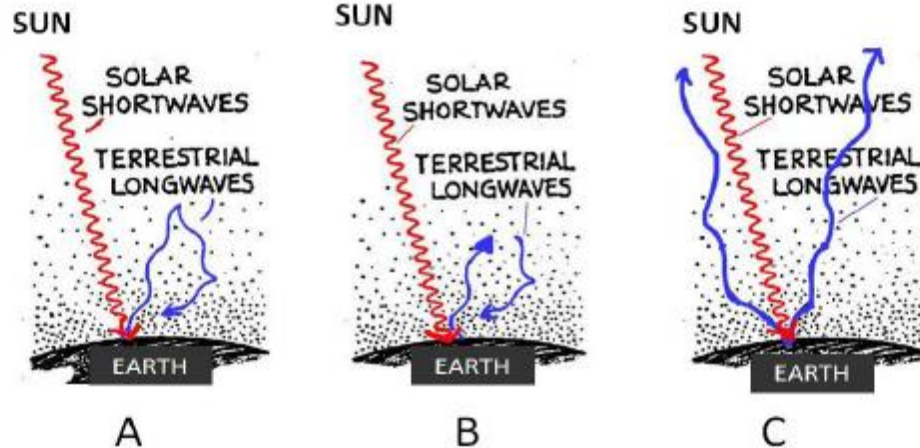
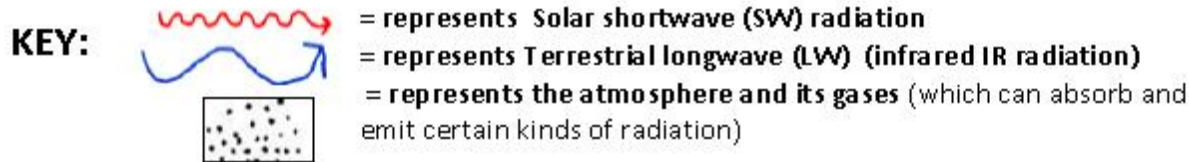
= represents the atmosphere and its gases (which can absorb and emit certain kinds of radiation)



5 minute GROUP WORK:
Answer Q7 – Q11

ANSWERS

SOLAR vs TERRESTRIAL RADIATION CLASS CONCEPTS SELF TEST



All 3 are illustrating ABSORPTION of incoming Solar SW by the EARTH's surface followed by outgoing RADIATION of LW Infrared from the EARTH's surface !

Q7. Which diagram above shows SW (solar radiation being reflected back to space?

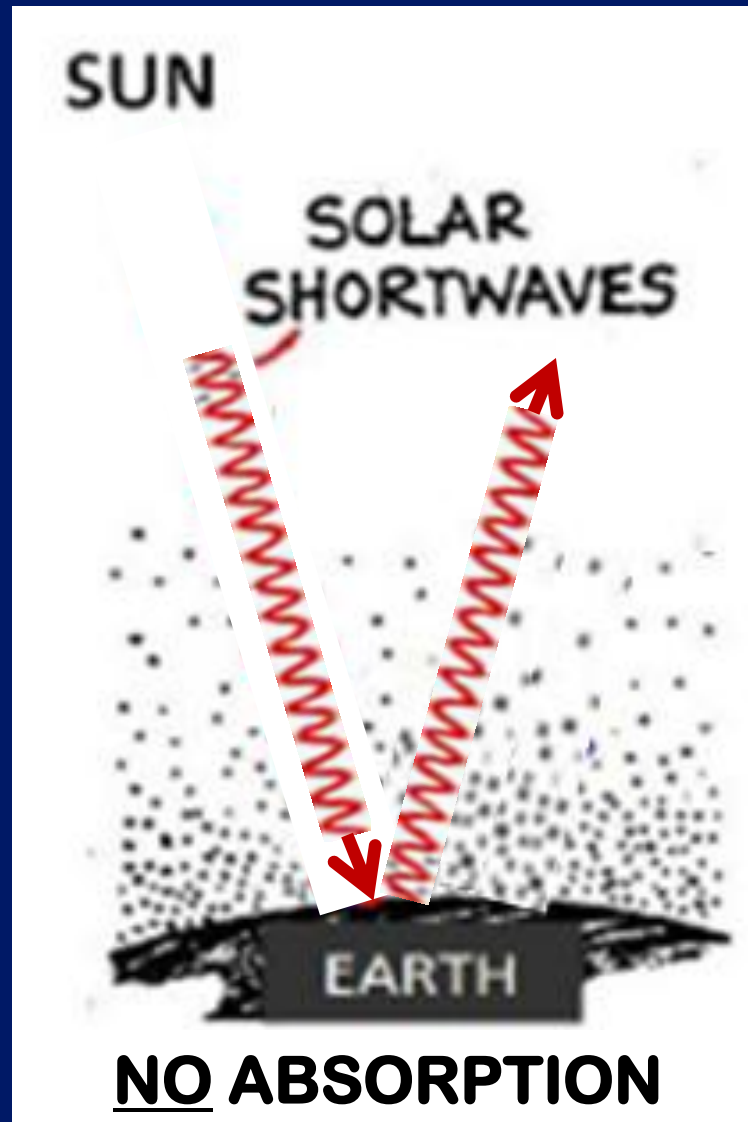
A

B

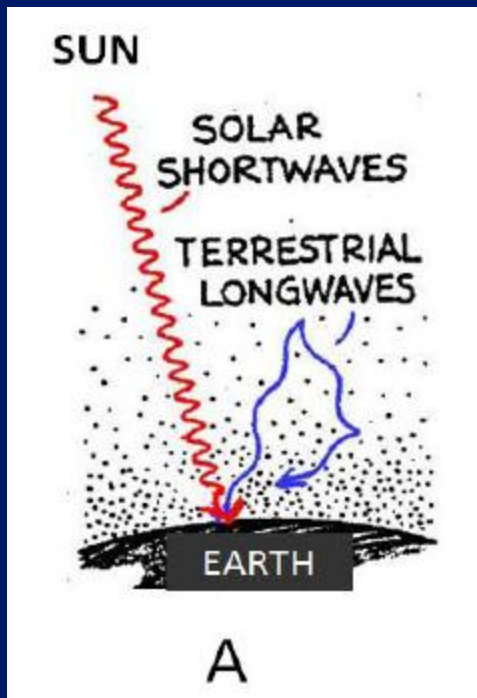
C

None of them

p 29



Here's the correct diagram to show **SW**
Solar being reflected back to space!



Q8. Diagram A shows LW (IR) terrestrial radiation “**bouncing off**” (or **reflecting**) the gases in the atmosphere and being sent back to Earth’s surface. (i.e. being reflected back to the surface by the gases without being absorbed by them.)

Is this an accurate depiction of how the Greenhouse Effect works?

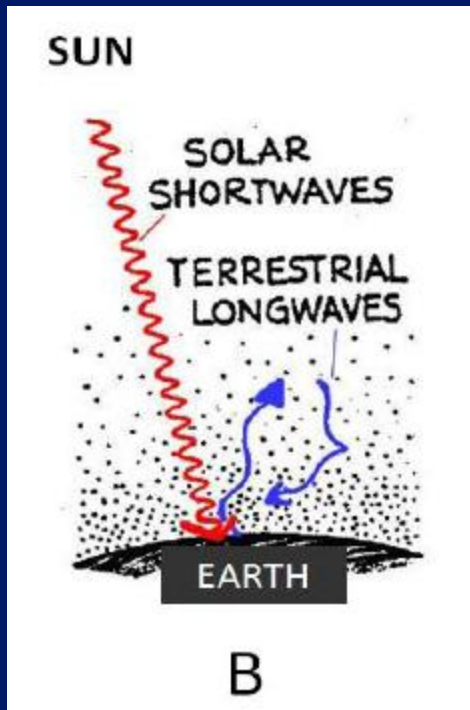
Yes

No

Why or Why not?

The LW (IR) radiation is **absorbed** by the GHG’s and then **re-radiated (or re-emitted)** back down to the Earth’s surface to warm it. The IR is **NOT reflected**.
IF IT’S REFLECTED IS NOT ABSORBED.

DON’T USE “BOUNCING or “REFLECTING” to describe the Greenhouse Effect process: **GH gases ABSORB & RE-RADIATE!**



Q9. Diagram B shows LW (IR) terrestrial radiation being absorbed and then emitted back down by the gases in the atmosphere.

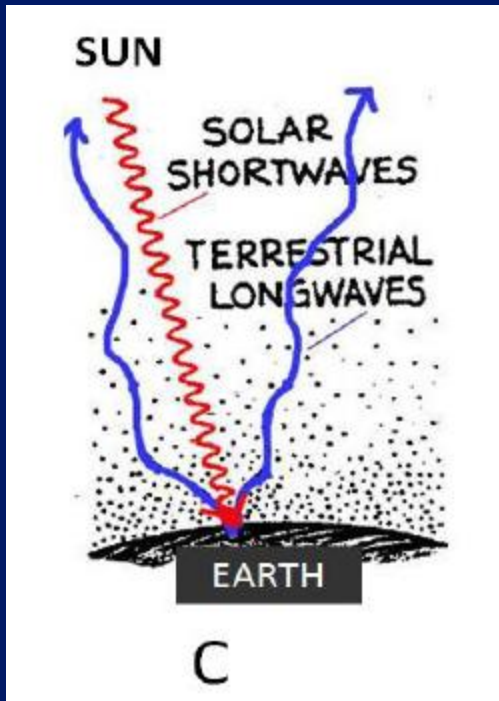
Is this an accurate depiction of how the Greenhouse Effect works?

Yes

Why or Why not?

(although it's not quite complete – more on this later)

LW (IR) being absorbed by GH Gases, and then emitted out again, (back down to the surface of the Earth) is **exactly how the GH Effect operates.**



Q10. Diagram C shows LW (IR) terrestrial radiation **going right through the atmosphere** out to space.

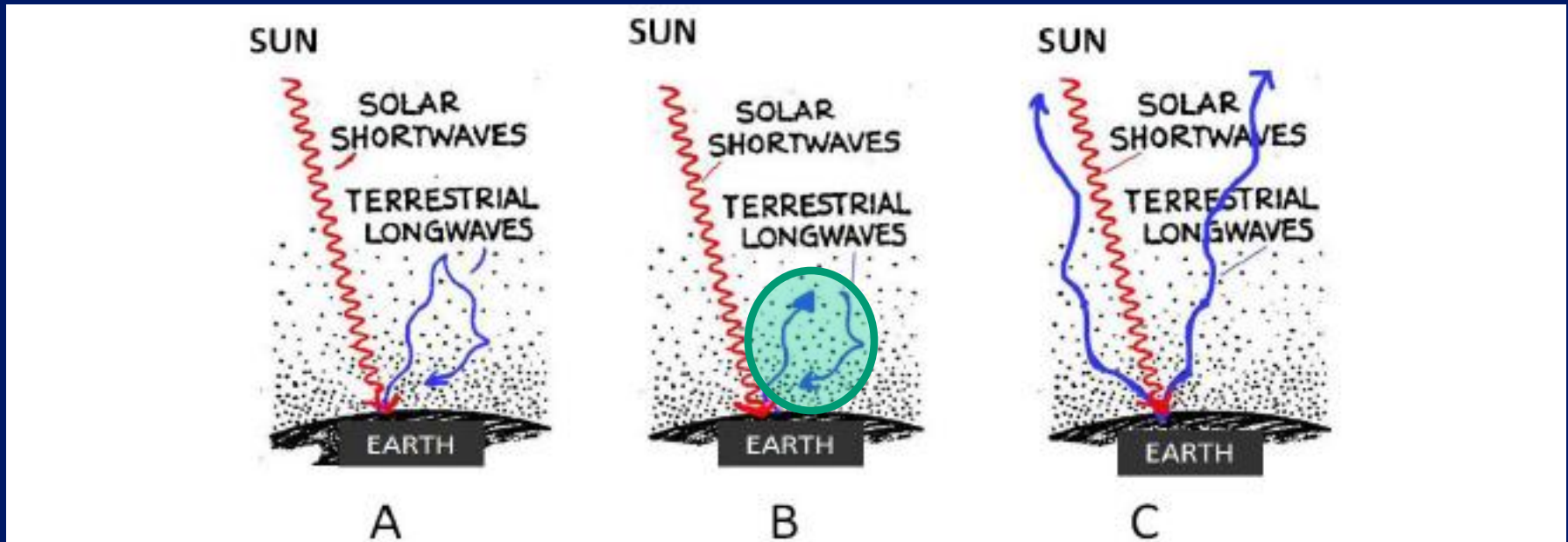
Is this an accurate depiction of how the Greenhouse Effect works?

Yes No Why or Why not?

Diagram C shows **ALL** the IR **leaving the Earth's surface** and **NOT** being absorbed at all !

If this happened the Earth's temperature would be below freezing!

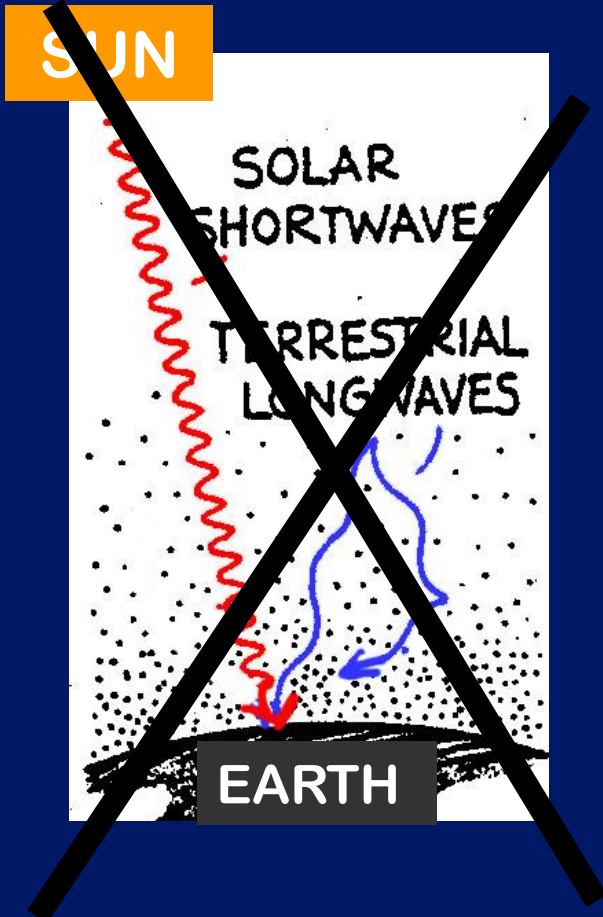
Q11. On the diagram that you think **best depicts the processes involved in the GREENHOUSE EFFECT**, CIRCLE the specific part of the diagram that represents the **Greenhouse Effect**:



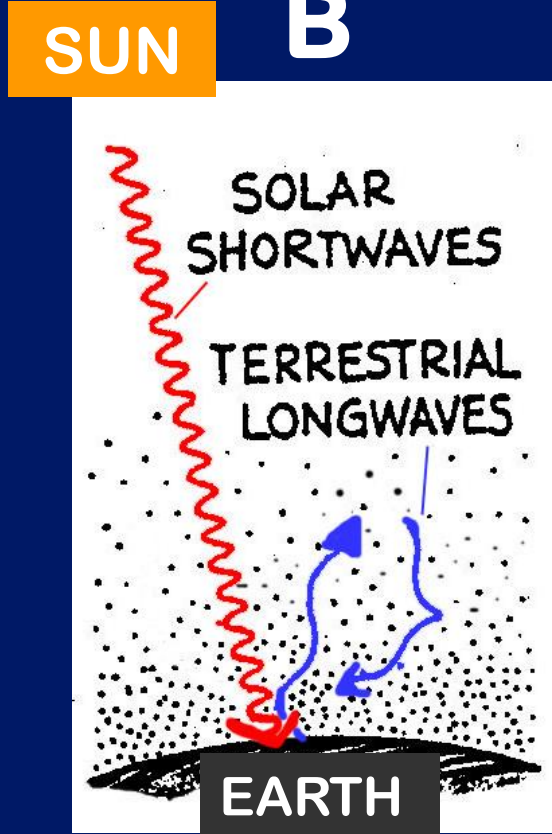
DO NOT CIRCLE any part of the **SUN's incoming SW!**
(Shortwave Radiation = IR + VIS)
SW is NOT part of the GH Effect!

Most accurate depiction of the GH Effect?

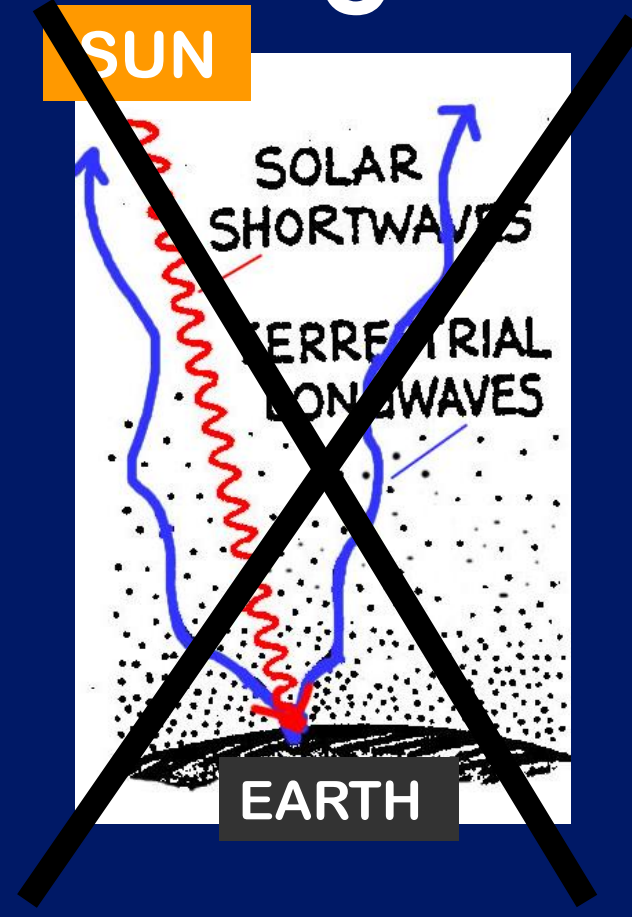
A



B



C



B is close, but actually, **NONE** of these is exactly correct . . .
for a better version:
see the bottom of p 29 in Class Notes.





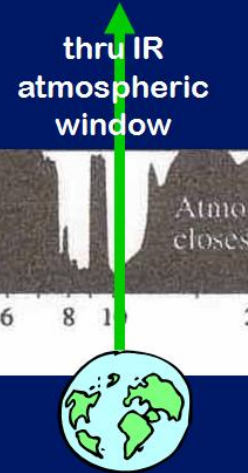
THINKING MORE DEEPLY . . .

① Some Incoming SW radiation from the SUN goes right through the atmosphere to Earth (w/o being absorbed)

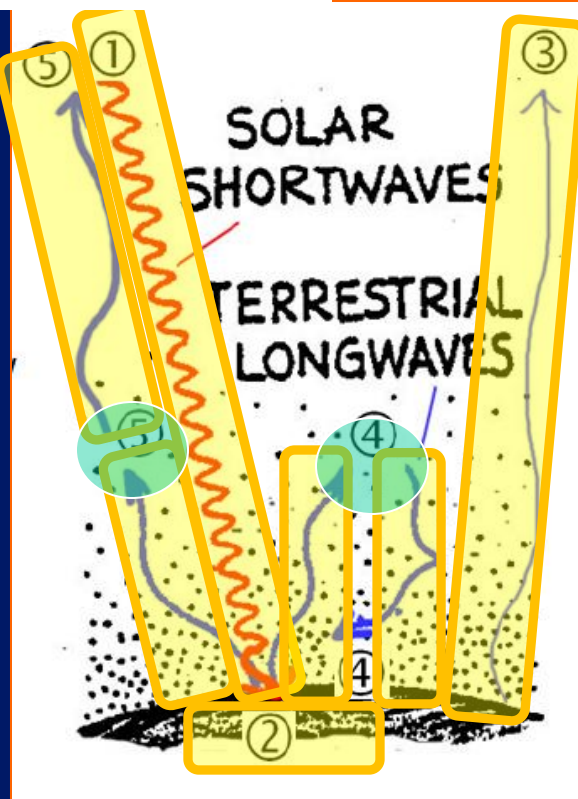
How? Through the UV-VIS Atmospheric Window!

③ Some IR radiation is emitted from the Earth's surface right out to space through "IR window"

Outgoing LW



⑤ Some IR radiation is **absorbed** by GH gases in the atmosphere, but is **emitted out to space** (not back to Earth)



④ Some IR radiation is **absorbed** by GH gases in the atmosphere and **emitted back to Earth**

② The Earth absorbs SW that reaches the surface

Absorption & re-emission by GH gases

Absorption & re-emission by GH gases

Bottom of p 29

HAPPY SEPTEMBER EQUINOX!

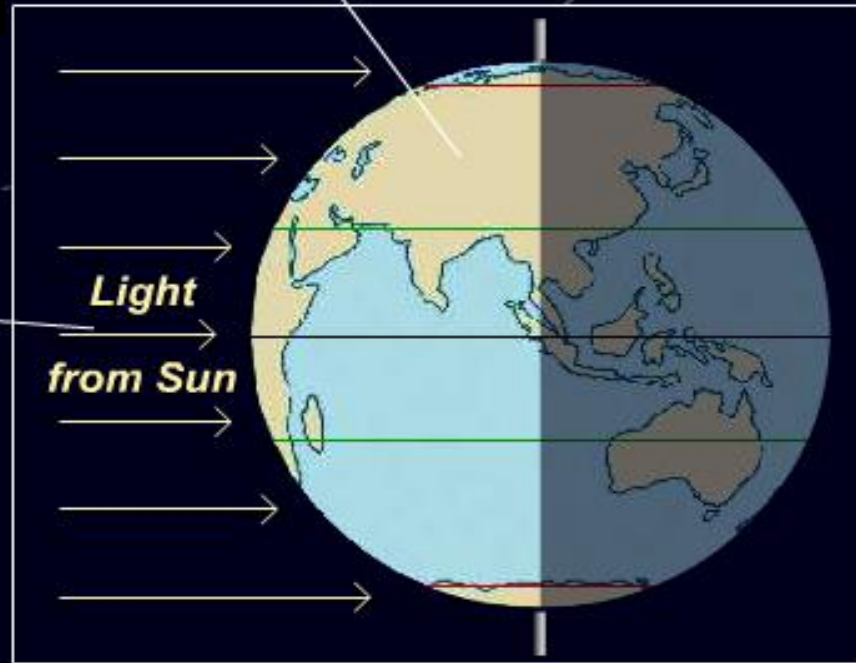
23 Sep 2015 @ at 1:22 AM Tucson Time

**Equinox =
“equal night”**

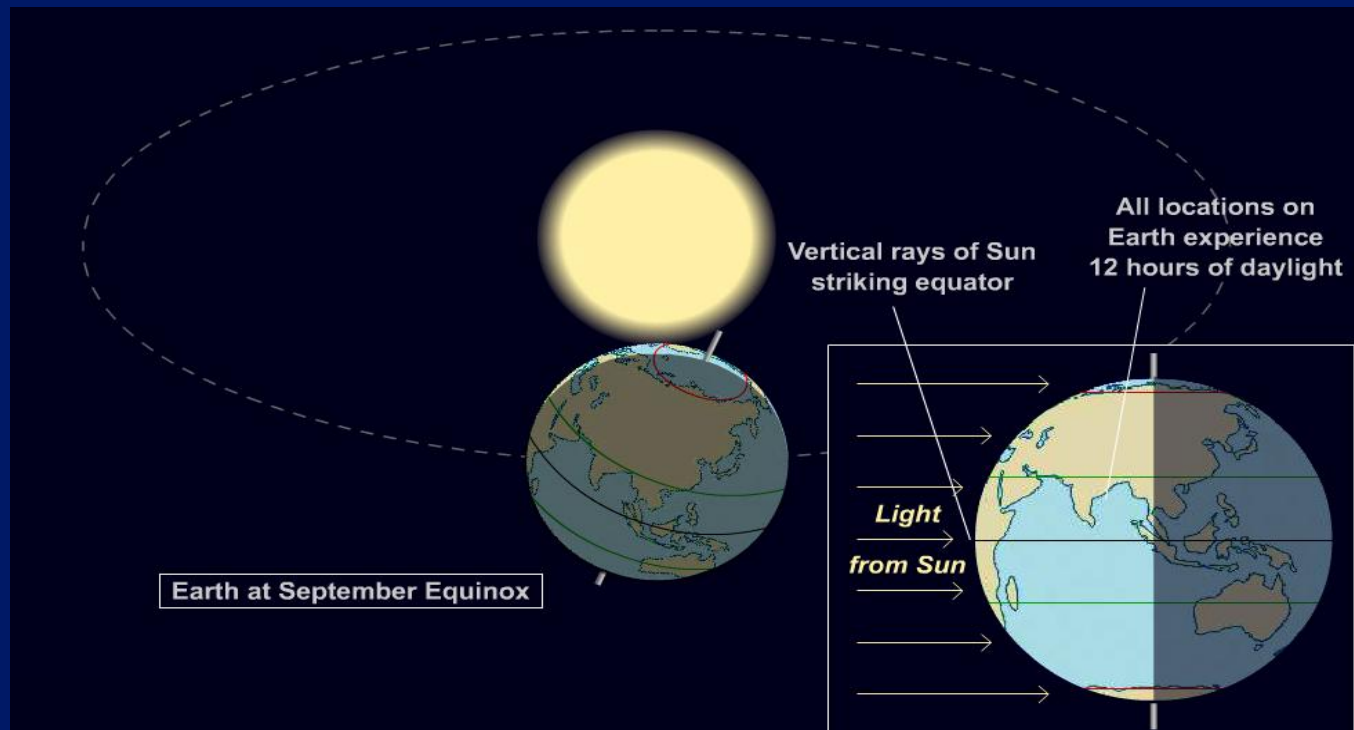
All locations on
Earth experience
12 hours of daylight

Vertical rays of Sun
striking equator

*Light
from Sun*

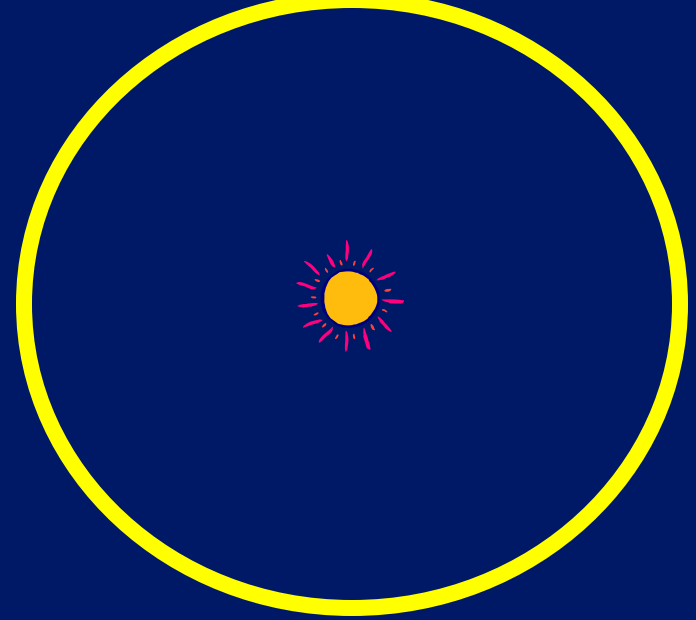


Today is the “Astronomical” “First Day of Fall” in the Northern Hemisphere!

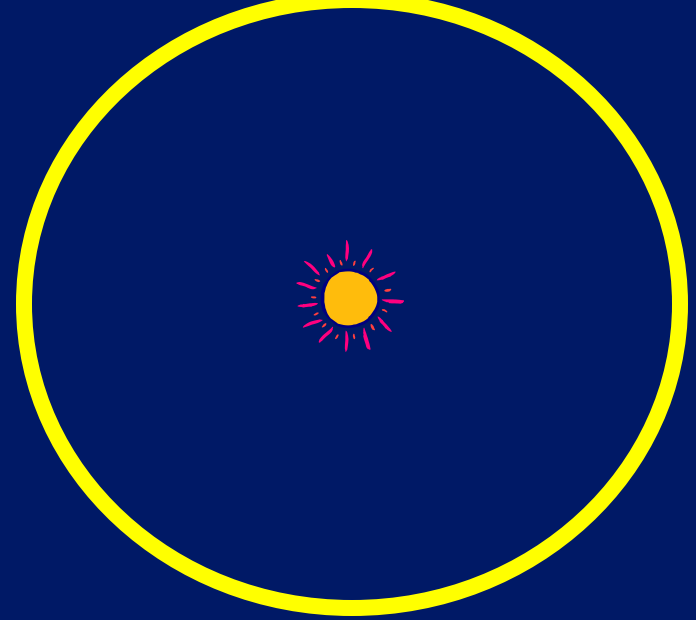


You can view the animation yourself at:

http://mesoscale.agron.iastate.edu/agron206/animations/01_EarthSun.html

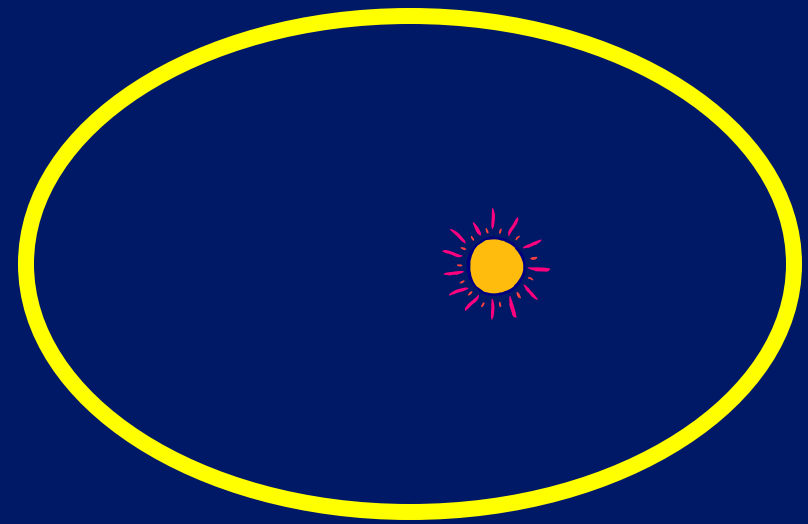


(1) Draw a **circular ORBIT** on the whiteboard & show **SUN's** location then use a **phone flashlight** to simulate Sun's radiation



(1) Draw a **circular ORBIT** on the whiteboard & show **SUN's** location then use a **phone flashlight** to simulate Sun's radiation

(2) Then, keeping the Earth's axis at a **constant 23 ½ angle**, and keeping the "north pole" axis 'tip' **always pointing to the "North Star"** – simulate a year of the Earth's Orbit – and discuss how the amount of radiation varies at different latitudes → seasonal changes!

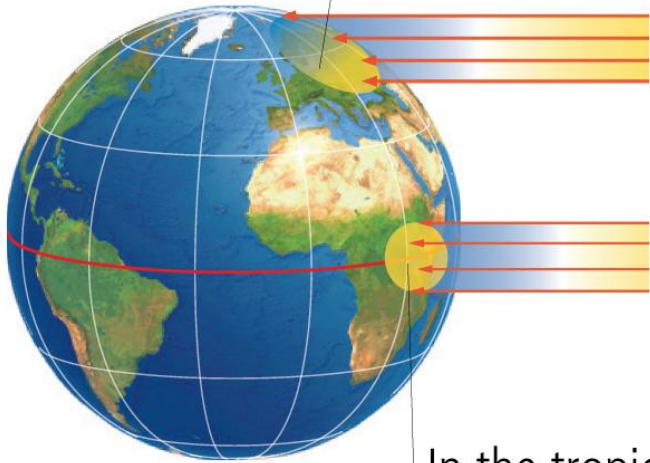


BUT . . . the Earth's orbit at this point in time is **NOT circular, it's **ELLIPTICAL** & the **SUN** is **NOT** in the center!**

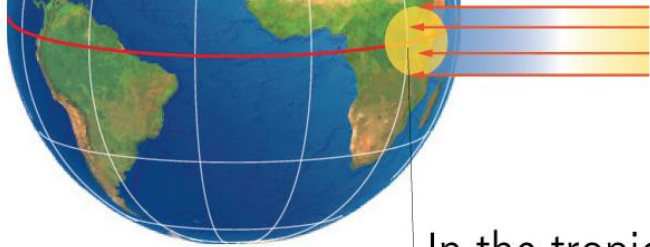
How does this change the amount of Solar radiation received in different hemispheres?

At what latitudes is this difference the most drastic?

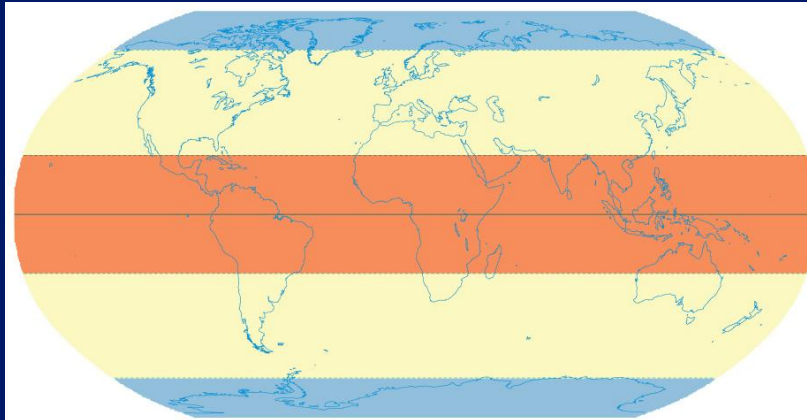
In the far north energy from the Sun is dispersed.



In the tropics energy from the Sun is concentrated.



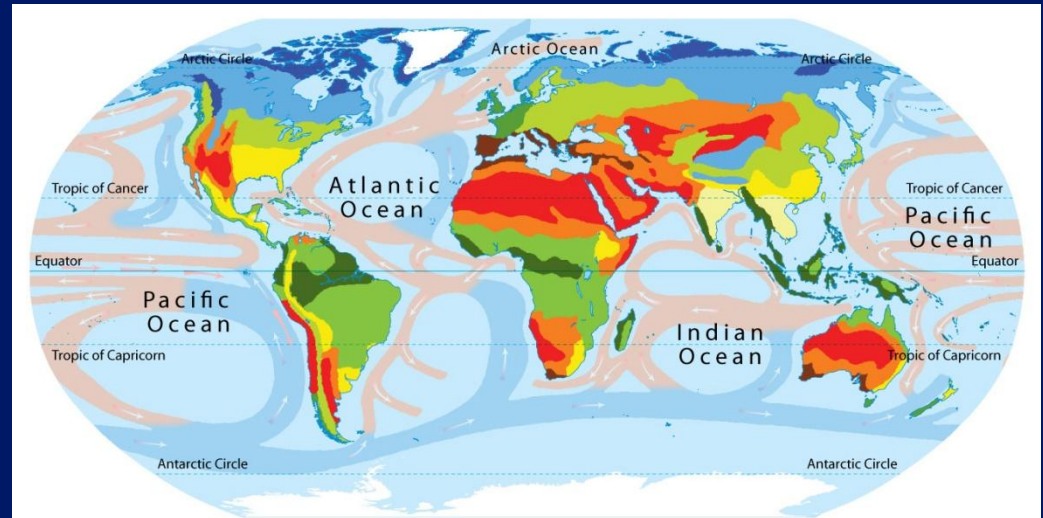
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Climatic bands

- Polar regions
- Temperate zones
- The tropics



Climatic zones

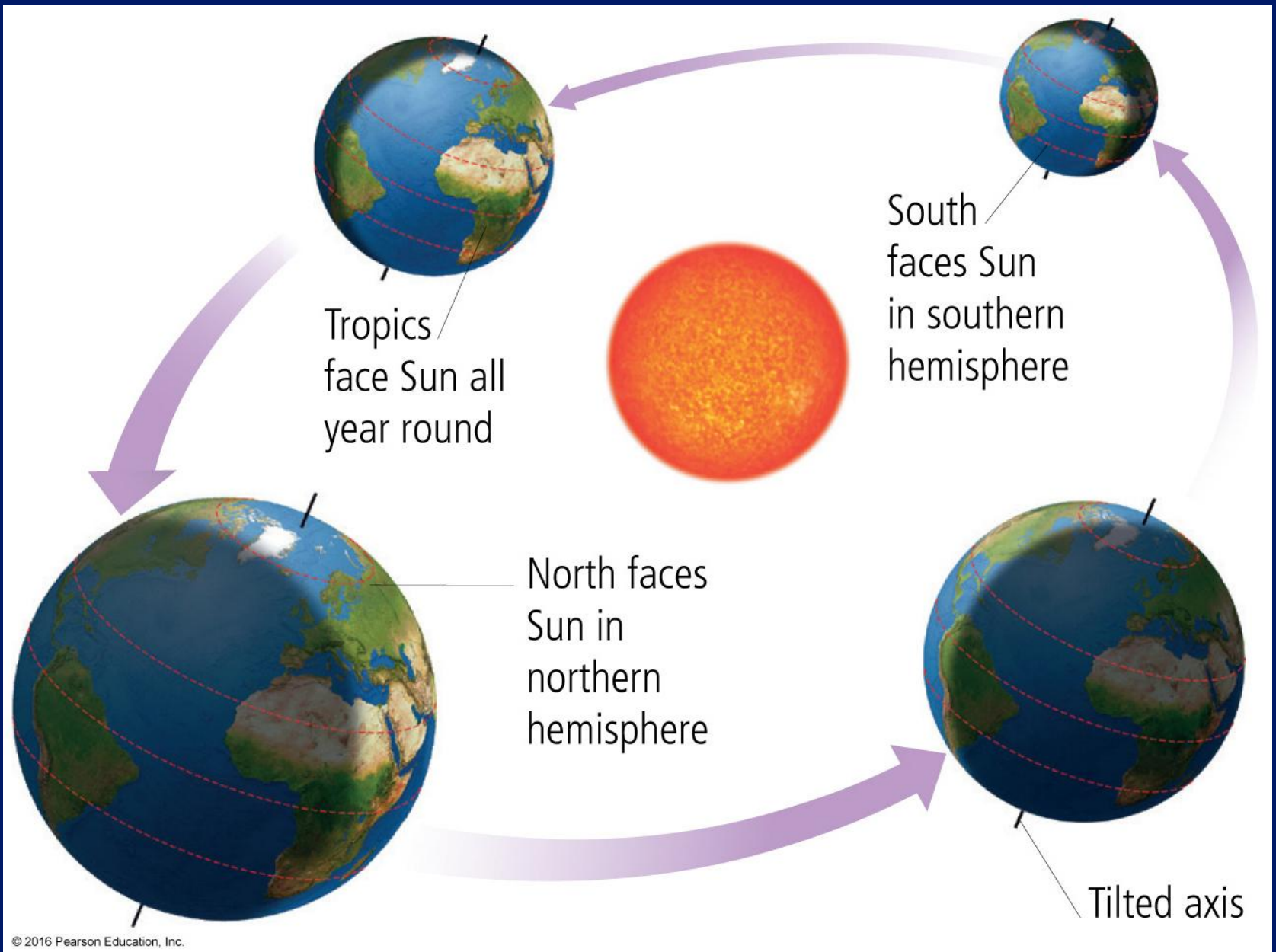
- | | | | | | |
|--|--|---|--|--|---|
| Ice cap | Tundra | Temperate | Mediterranean | Arid | Humid equatorial |
| Subarctic | Continental | Warm temperate | Semi-arid | Hot humid | Tropical |

Ocean currents

- Warm
- Cold

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READ: Dire Predictions pp 10-15



READ: Dire Predictions pp 10-15

We'll address more on

EARTH-SUN relationships ,

how they've changed in the past,

and how this is an important

NATURAL cause of

CLIMATE CHANGE

in Topics #10 & #11



CHECKPOINT

THINK for 15 seconds

TABLE CHAT for 15 seconds

What's your most burning question about:

EARTH-SUN RELATIONSHIPS ,

the SEASONS

and GLOBAL CLIMATE PATTERNS?

**Topic # 6 ATMOSPHERIC
STRUCTURE
&
CHEMICAL COMPOSITION**

**All about the GASES IN THE
ATMOSPHERE, esp.
GREENHOUSE GASES!**

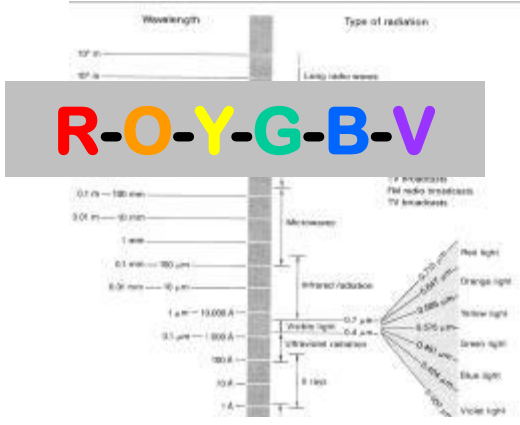
OBJECTIVES:

To understand:

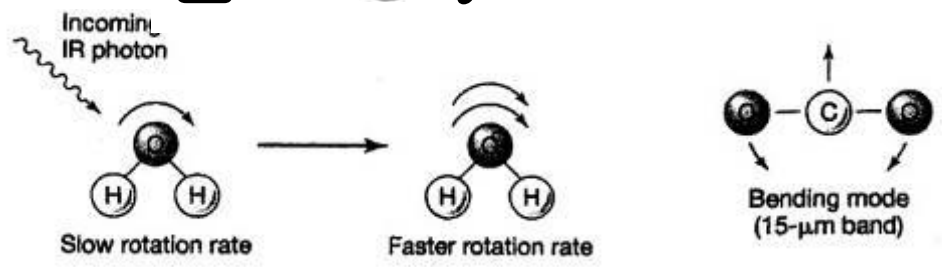
- the **VERTICALSTRUCTURE** of the atmosphere & its relationship to temperature
- which **GASES** are in the atmosphere
- **where** they are concentrated, and
- why gases at different levels are linked to the **Greenhouse Effect** and also **Ozone Depletion**
(which are NOT the same thing!!!)



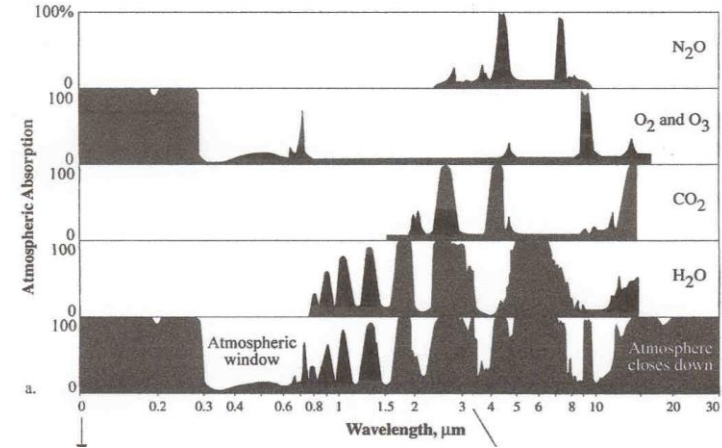
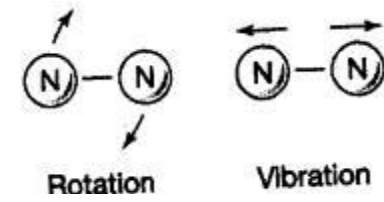
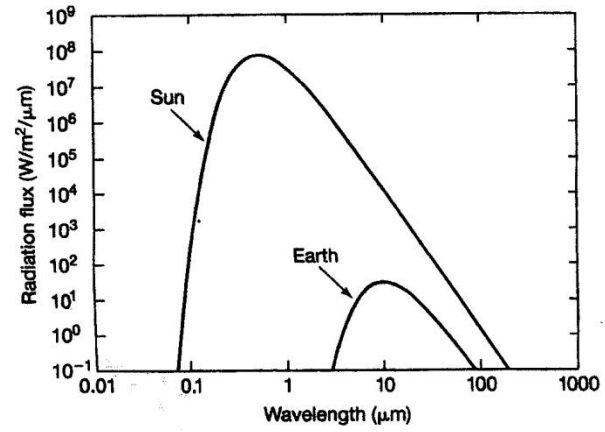
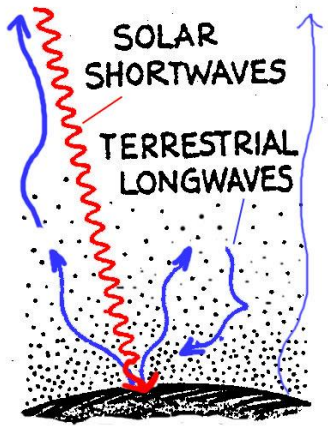
Things you've seen before that will all come together under this topic:



$$E = \sigma T^4$$



$$E = hc / \lambda$$



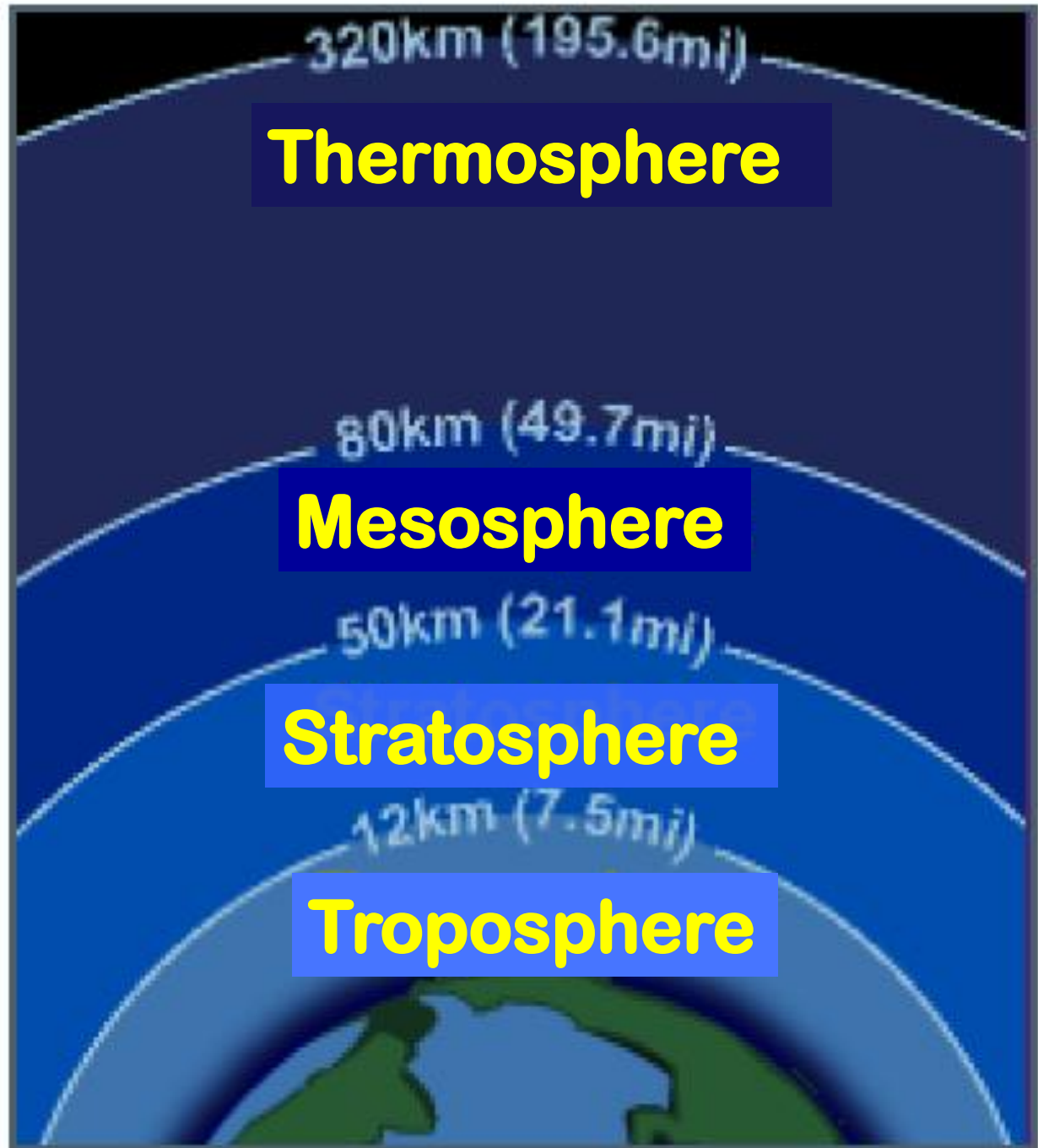
$$\lambda_m = a / T$$

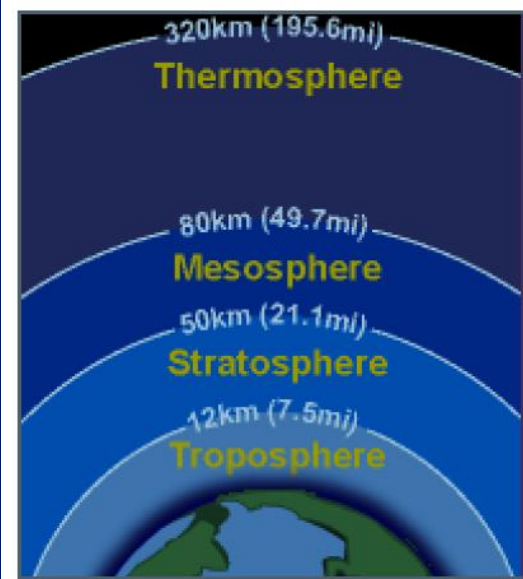
**We travel together, passengers in a
little space-ship, dependent on its
vulnerable supplies of air and soil.**

~ Adlai Stevenson



The atmosphere has a “structure” of different named layers :





Mesopause

Boundary between the mesosphere and the thermosphere.

Stratopause

Boundary between the stratosphere and the mesosphere.

Atmospheric ozone layer

Layer within the stratosphere. Absorbs ultraviolet solar radiation, so warming the surrounding atmosphere.

Tropopause

Boundary between the troposphere and the stratosphere.

Exosphere

Outermost layer of the atmosphere. Extends to about 10,000 km (6,000 miles).

Thermosphere

Extends to about 640 km (400 miles).

Mesosphere

Rises from about 50 to 80 km (30 to 50 miles) above the surface. Air becomes cooler as the altitude increases.

Stratosphere

Extends upward to a height of about 50 km (30 miles). Contains atmospheric ozone layer. Temperature increases with altitude through the stratosphere, inhibiting vertical air currents, and making the stratosphere highly stable, in contrast to the troposphere.

Troposphere

Layer in contact with Earth's surface. Extends from the surface to about 8 to 17 km (5 to 10 miles). Air temperature decreases with altitude, leading to instability. Less dense air sits below more dense air, which results in air movements and storm generation. "Weather" takes place almost exclusively within the troposphere.

Sea level

MAKE A GRAPH:



ALTITUDE
Kilometers (km)
above sea level)

100

0

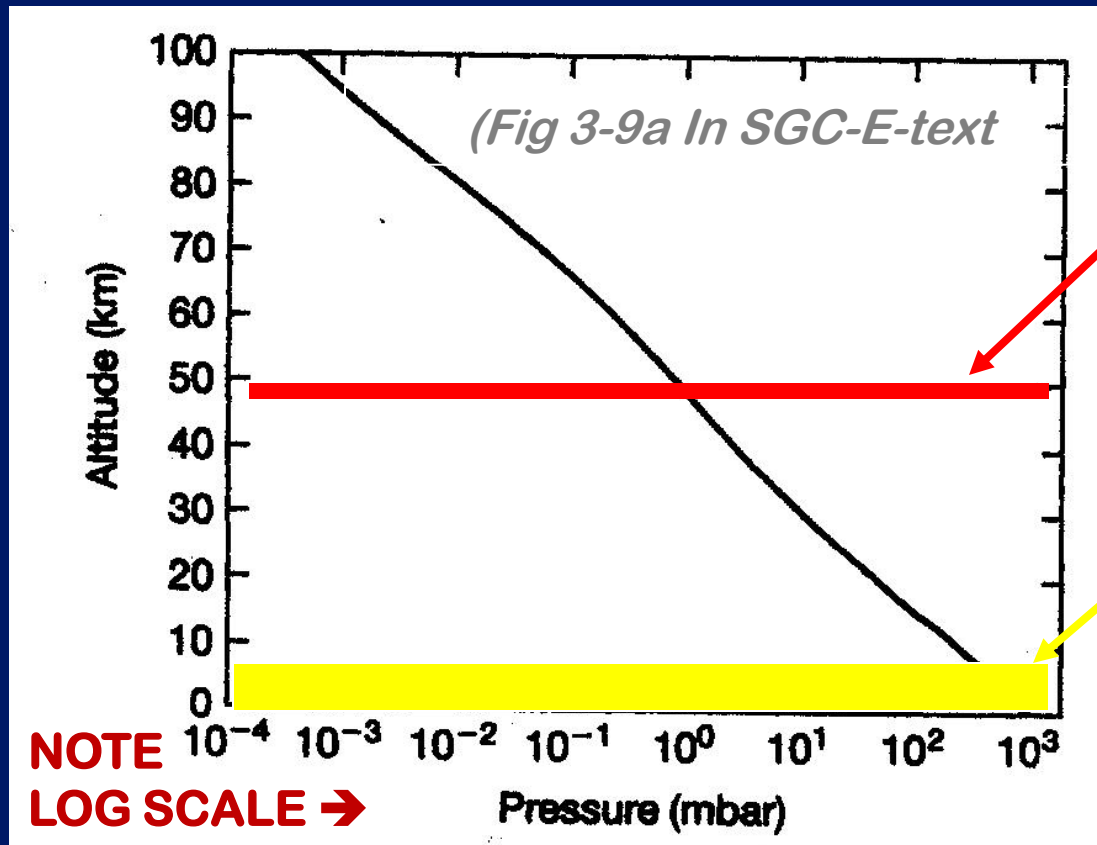
Low

ATMOSPHERIC PRESSURE

High

Atmospheric Pressure = weight of the air column above

Atmospheric Pressure & Mass Vary with Height



**99% of mass
lies below ~
50 km (top of
Stratosphere)**

**50% of mass
lies below ~ 6
km (middle
Troposphere)**

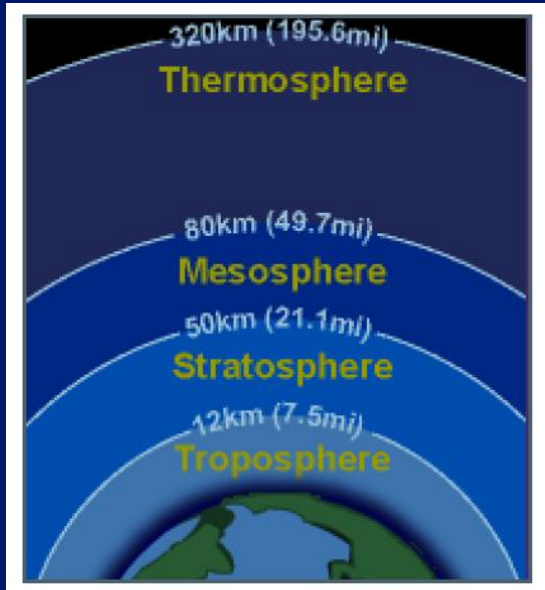


MAKE A GRAPH:

100



ALTITUDE
Kilometers (km)
above sea level)



0

Thermosphere 80 km – 320 km

Mesosphere 50 km – 80 km

Stratosphere 12 km – 50 km

Troposphere 0 km – 12 km

Now erase pressure and
sketch in the layers on
your graph with a dashed
line - - - - -

MAKE A GRAPH:



ALTITUDE
Kilometers (km)
above sea level)

Sketch in
Temperature
curve

100

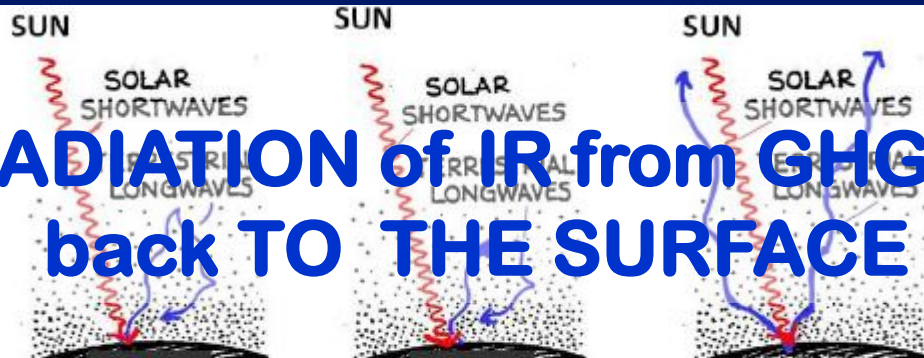
0

COLD

ATMOSPHERIC TEMPERATURE HOT

The ATMOSPHERE IS PRIMARILY HEATED FROM BELOW (by IR energy radiated up from the EARTH'S SURFACE)

**RADIATION of IR from GHG's
back TO THE SURFACE**



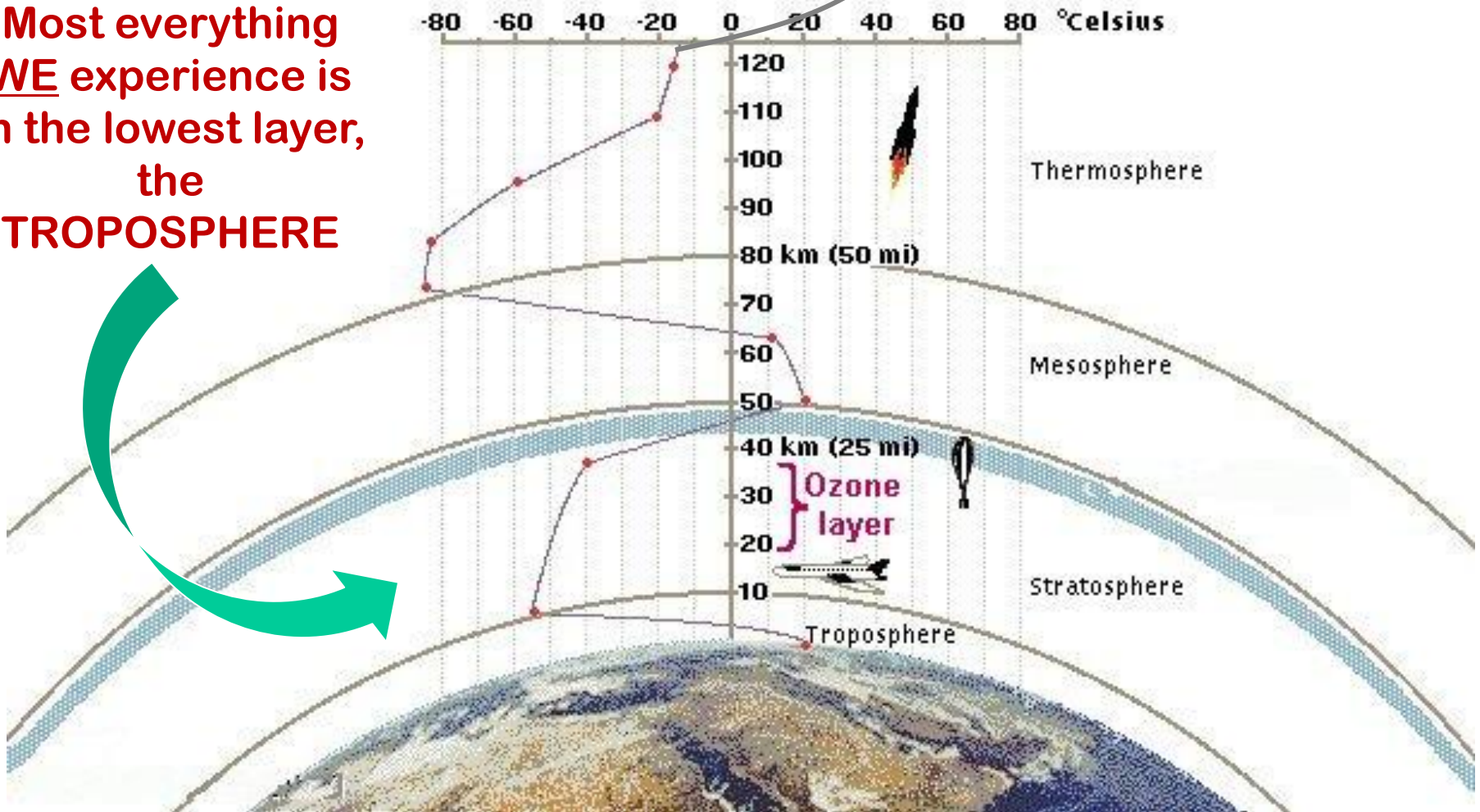
ABSORPTION of ↓ Solar SW

RADIATION of ↑ IR from surface

These layers have different thicknesses and temperatures...

Most everything WE experience is in the lowest layer, the **TROPOSPHERE**

This zig-zag line is showing changes in temperature with altitude



The Vertical Structure of the Atmosphere

KEY CONCEPT:

The atmosphere's vertical structure is defined by **CHANGES** in the trend of **TEMPERATURE** with height.

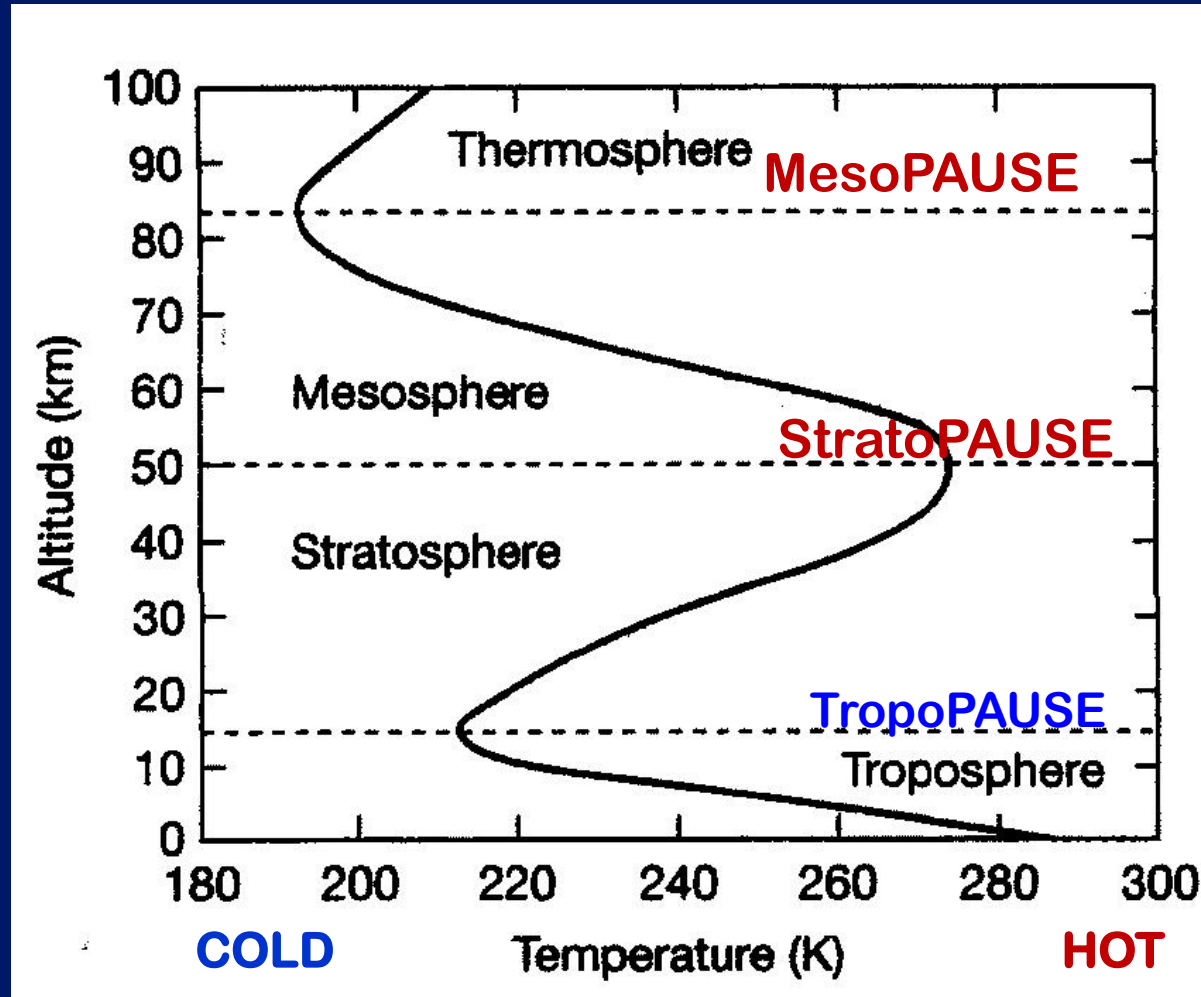
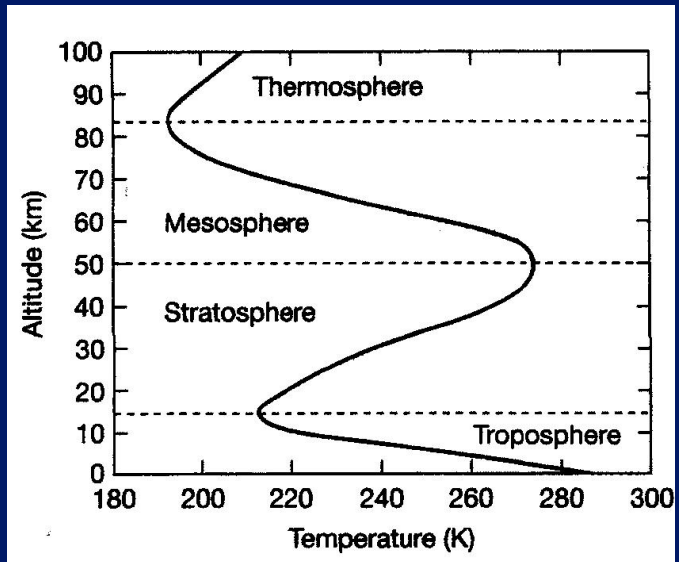


Figure 3-9b in SGC E-text



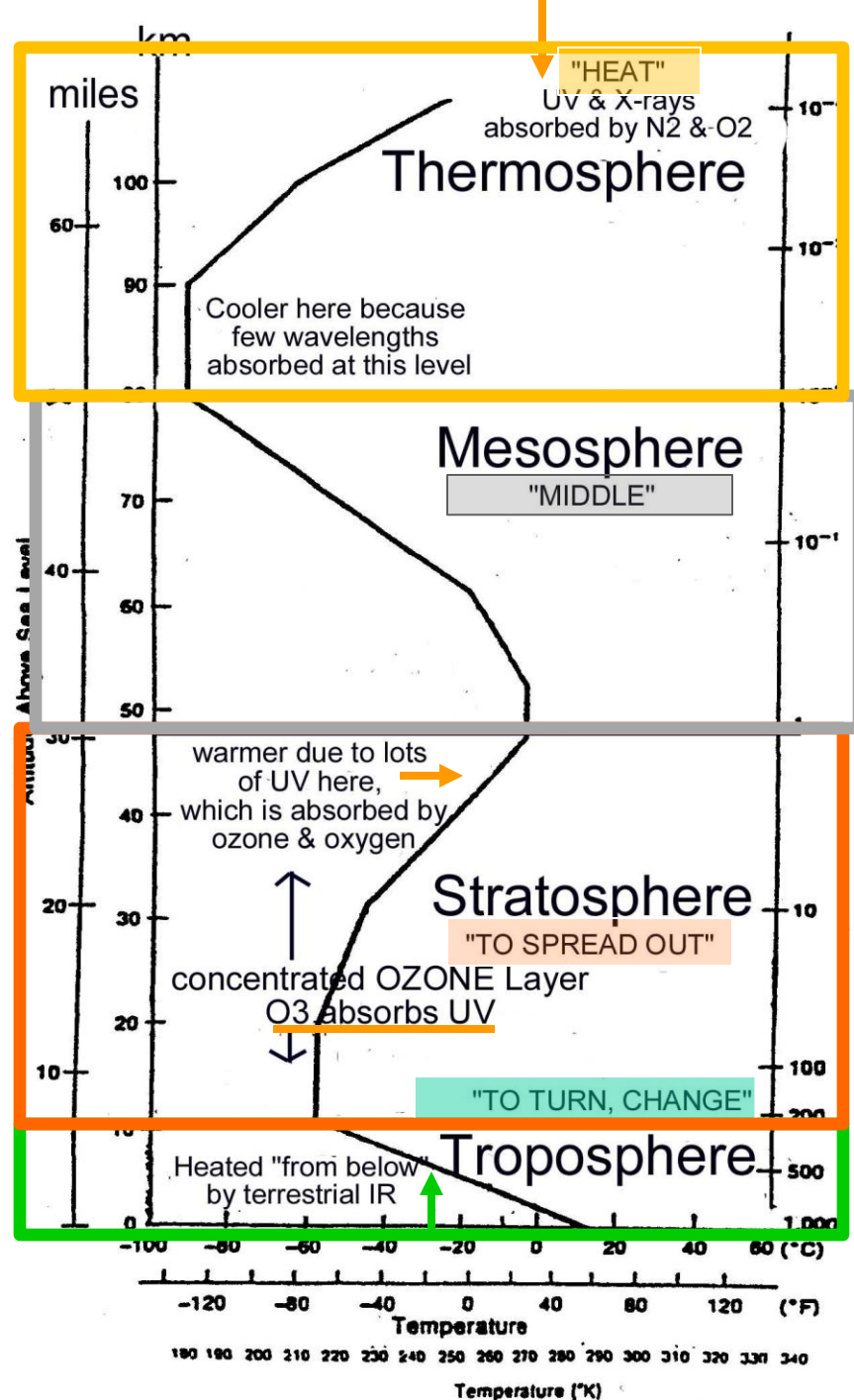
The changes in temperature with height are the result of:

differential absorption of shortwave (SW) & longwave (LW) radiation

by atmospheric GASES concentrated at various altitudes.

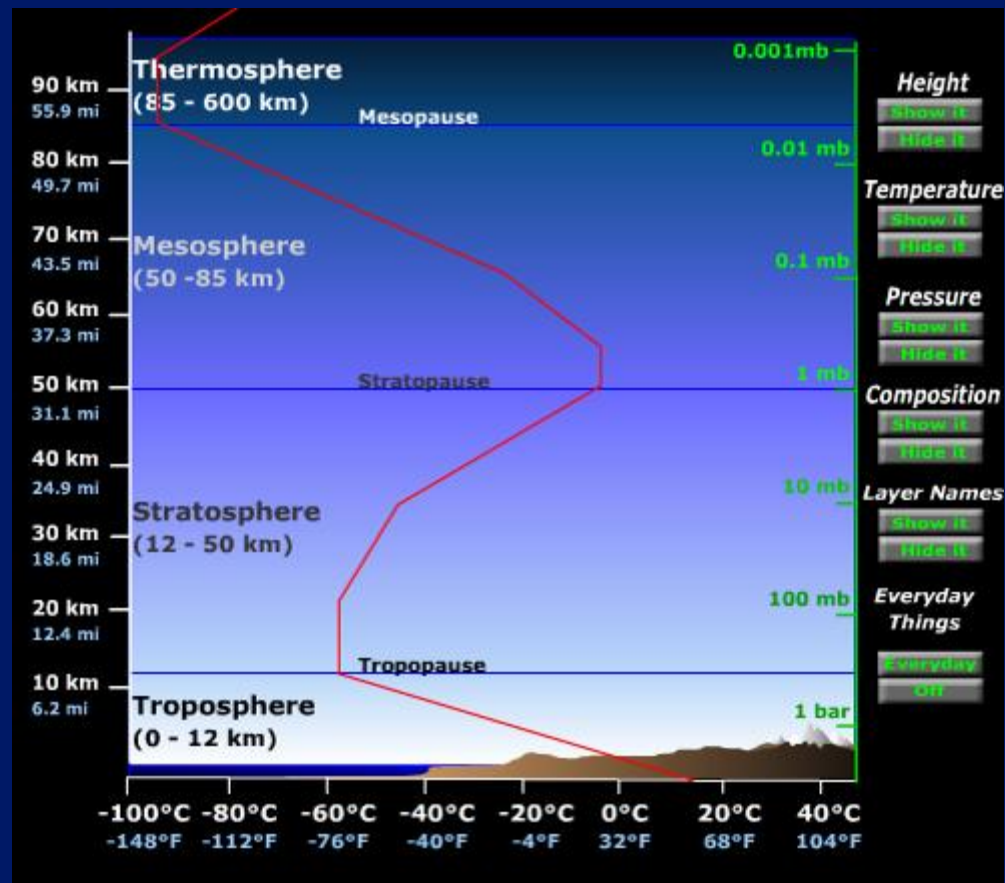
SUMMARY:
 Here's why
 these
 changes in
 temperature
 occur :

Start at the
SURFACE →



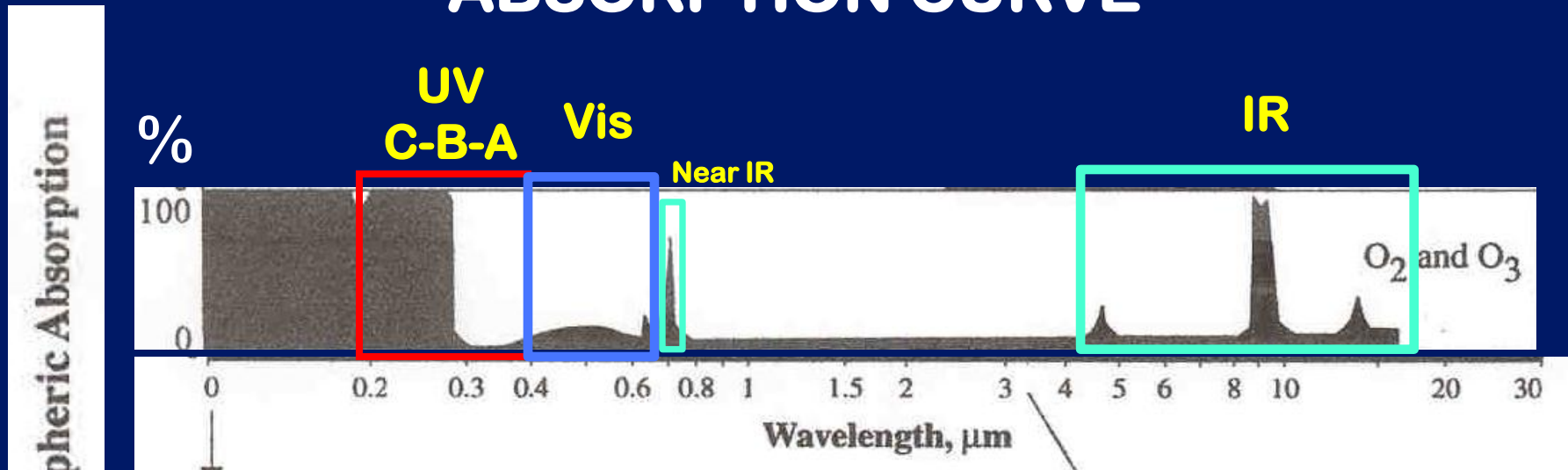
A nice online review . . .

<http://earthguide.ucsd.edu/earthguide/diagrams/atmosphere/index.html>



REVIEW: The pattern of electromagnetic wavelengths that are **absorbed & emitted** by a particular atom (or combination of atoms)

is called its **ABSORPTION SPECTRUM** or its **ABSORPTION CURVE**



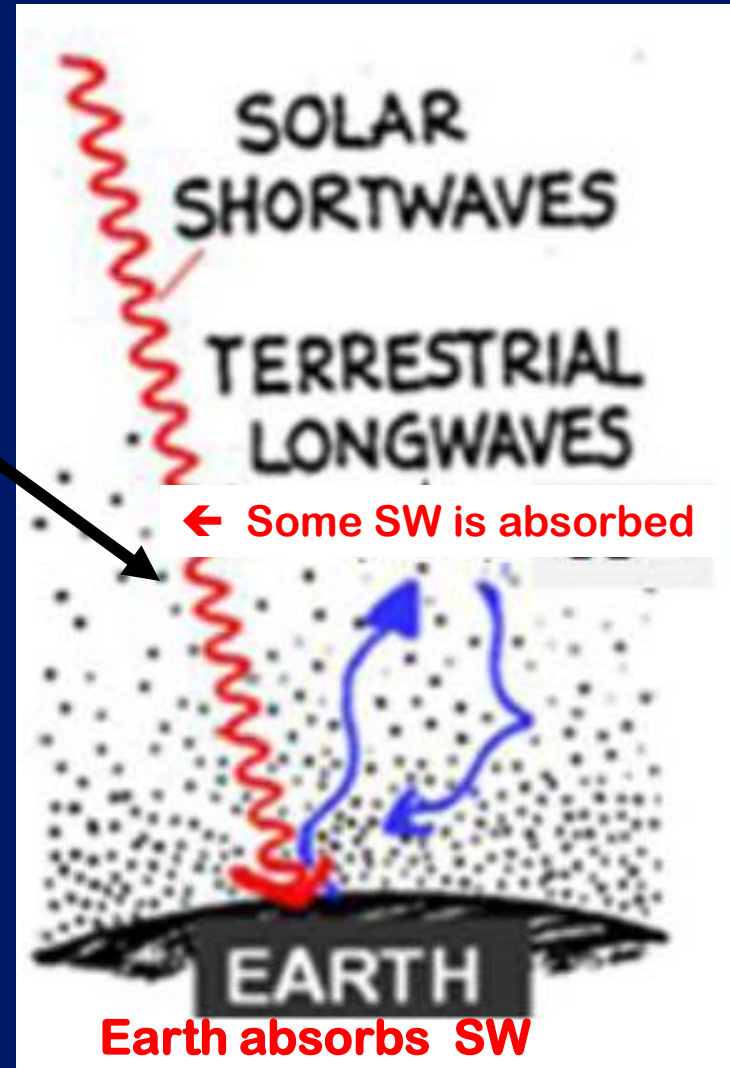
**The Absorption curve for
Ozone / Oxygen**



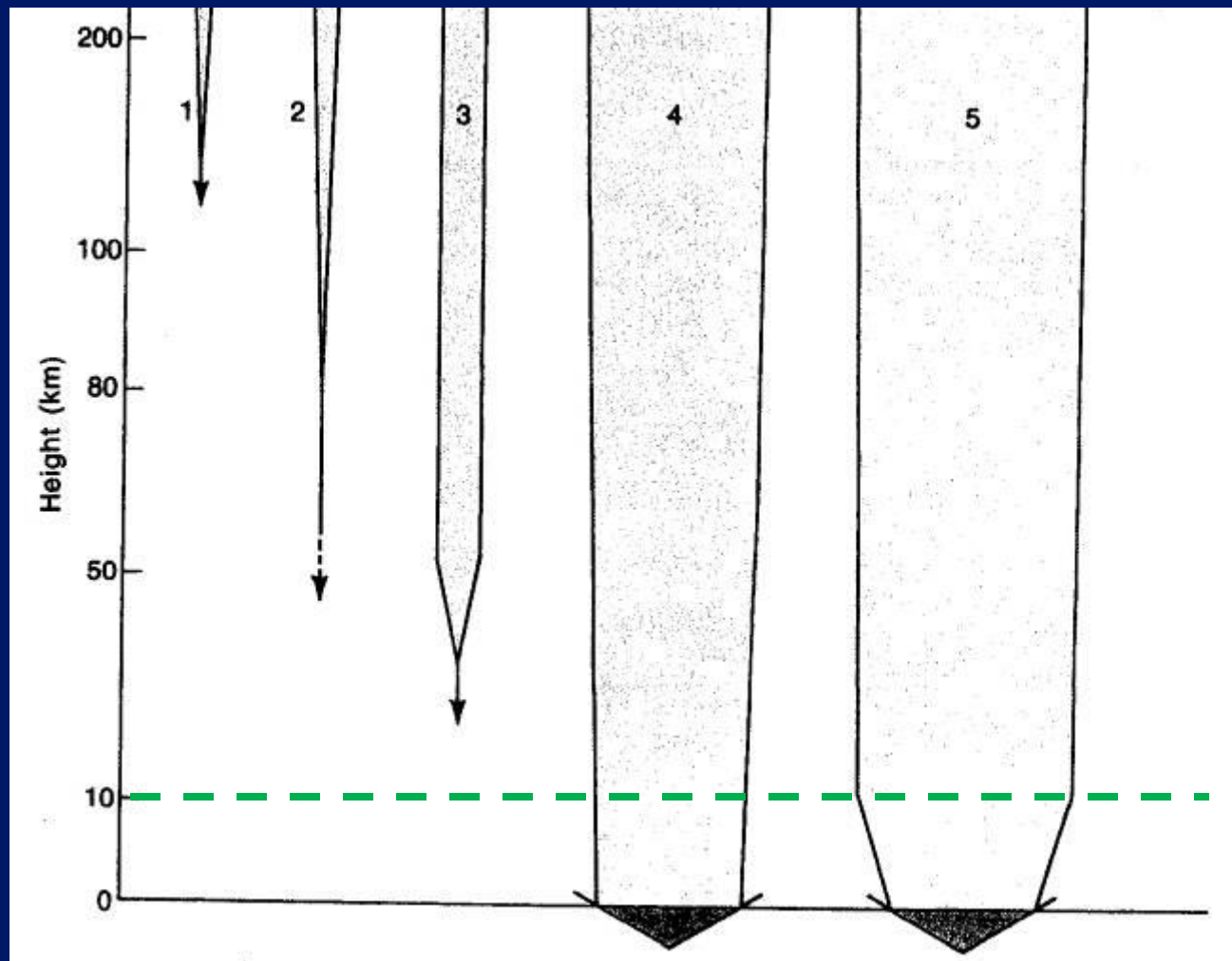
There's one more thing to correct in our the depiction of incoming Solar

Some SW radiation
gets absorbed on
its way down to the
surface!

(in addition to
terrestrial LW (IR)
radiation being
absorbed in the GHE)



How incoming **SOLAR** radiation of different wavelengths gets **TRANSMITTED** or **ABSORBED** by different gases on its way to the Earth's surface



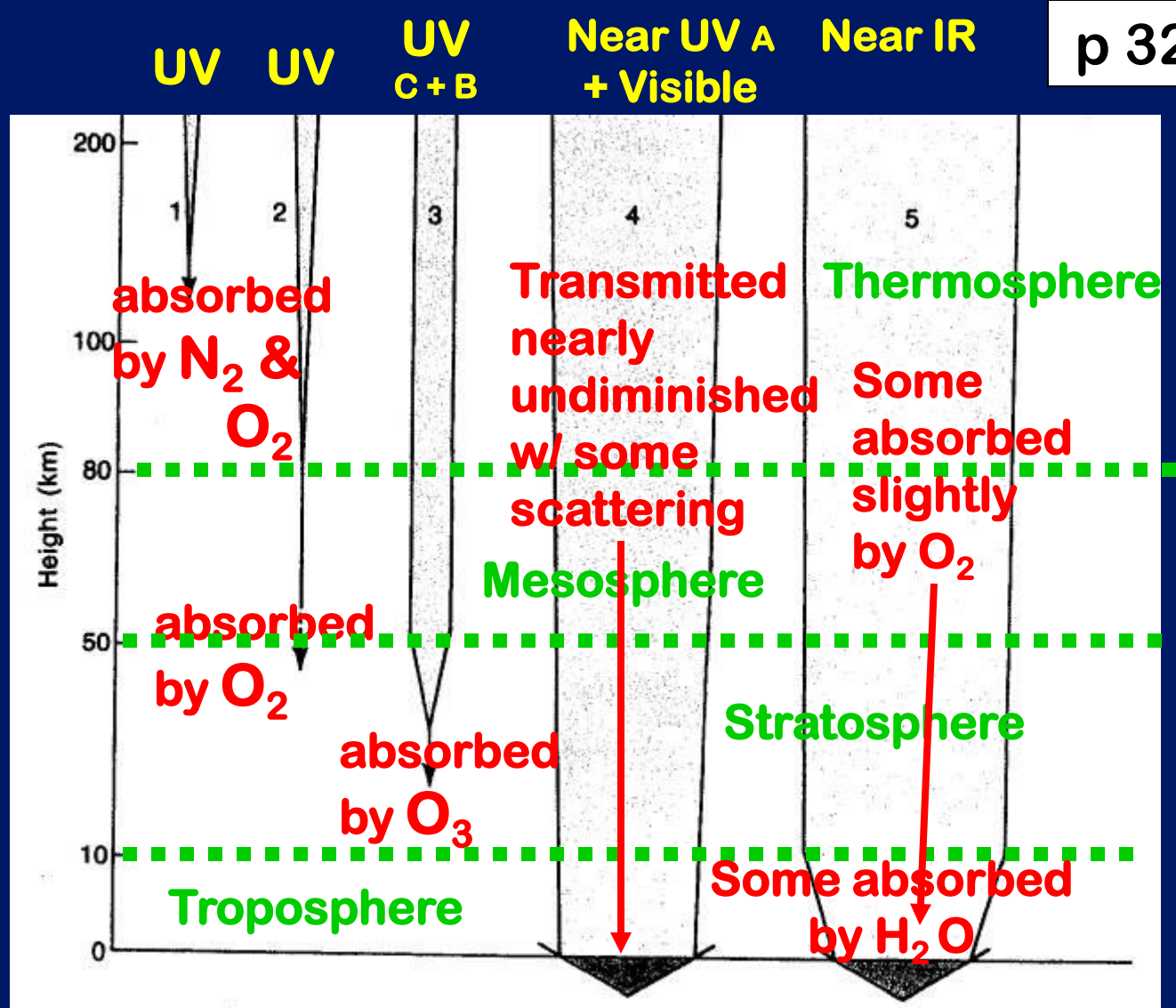
1. UV, $\lambda < 0.12 \mu\text{m}$, absorbed by N_2 and O_2 in upper atmosphere
2. UV, $0.12 \mu\text{m} \leq \lambda < 0.18 \mu\text{m}$ absorbed by O_2
3. UV, $0.18 \mu\text{m} \leq \lambda < 0.34 \mu\text{m}$ absorbed by O_3 in ozone layer
4. Near UV and visible, $0.34 \mu\text{m} \leq \lambda < 0.7 \mu\text{m}$ transmitted nearly undiminished except for scattering
5. Near IR, $0.7 \mu\text{m} \leq \lambda < 3.0 \mu\text{m}$, absorbed slightly by O_2 and in troposphere by H_2O

Reminder: *Ultraviolet radiation:* UVC = 0.20 - 0.29 UVB = 0.29 - 0.32 UVA = 0.32 - 0.40 μm

UV rays < .32 μm
 very harmful to
 life on Earth arrows
 1, 2 + 3



How incoming
SOLAR
 radiation of
 different
 wavelengths
 gets
TRANSMITTED
 or **ABSORBED**
 by different
 gases
 on its way to
 the Earth's
 surface



1. UV, $\lambda < 0.12 \mu\text{m}$, absorbed by N_2 and O_2 in upper atmosphere
2. UV, $0.12 \mu\text{m} \leq \lambda < 0.18 \mu\text{m}$ absorbed by O_2
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5. Near IR, $0.7 \mu\text{m} \leq \lambda < 3.0 \mu\text{m}$, absorbed slightly by O_2 and in troposphere by H_2O



CHECKPOINT

THINK for 15 seconds

TABLE CHAT for 15 seconds

What's your most burning question?

ATMOSPHERIC COMPOSITION

Which gases?

What concentration?

Which ones are

Greenhouse Gases (GHG)?

Where do the GHG's come from?

**Which GHG's are changing in
concentration due to**

HUMAN ACTIVITIES?



**OVER THE WEEKEND
STUDY THE TABLES ON**

Pages 33 & 34

**to familiarize yourself with each
of the GHG's**

Then get ready for the

“NAME THAT GAS!”

Team Competition on Monday!

THINKING DEEPLY WRAP-UP



Participation Point Activity:

Get a blank index card from you group folder,
put Name & Group # on it

Pick one of the following,
reflect on what you learned today about it,
& explain this in a short paragraph →

What we did today . . .

- 1) Reviewed the **LEARNING PHILOSOPHY** of the course
- 2) Dug deeper into the pathways of **incoming solar and outgoing terrestrial radiation** and how the **Greenhouse Effect** is involved
- 3) Learned about **Earth-Sun Relationships** and what the an EQUINOX is – and how this causes the seasons and determines Earth's climate regions
- 4) Sketched the **Vertical Structure of the Atmosphere** and learned what causes it – including the role of certain gases in different layers

What to do: on the index card w/ name + Group #

Pick one of the above , **reflect on what you learned today about it**, & explain this in a short paragraph

SEE YOU MONDAY



GO CATS!